

Normal Mode Analysis with the Chesapeake Bay

Under Non-Optimal Conditions

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Normal Mode Analysis (NMA)

- Chesapeake Bay focused
- Calculate eigenmodes using a Helmholtz decomposition
- Use the eigenmodes as an orthonormal basis set
- Decompose a data set (velocity fields) into their modal components (amplitudes)
- Produce a time-dependent power spectrum
- Make predictions using eigenmodes over full domain

Previous Work

- Zel'Dovich - 1985 - vector decomposition
- Eremeev, Kirwan, et al. - 1992 - Black Sea
- Lipphardt, Kirwan, et al. - 2000 - Monterey Bay

NMA - Difficulties

- Computing eigenmodes over complex boundaries can be questionable (how close to ortho-normal).
- NMA requires full data sets (in space).
- Hardware sensors collect data sporadically (time irregular).
- Eigenmodes exhibit two behaviors (global and local).
- Differing methods exist to extract amplitudes (Galerkin, LSHA, other fitting).

NMA – Non-Optimal Conditions

- Chesapeake Bay – less than perfect
 - Large estuary – 180x50 miles
 - Significant sources – Atlantic Ocean, large watershed, five major rivers
 - Boundary – fractal
 - Geometric variation – many regions of varying sizes
 - Man-made sources – several major cities, nuclear reactor
 - Salinity variation – southern – salty, northern – fresh
 - Shallow water bathymetry – average depth 8.4m, max ~ 30m
 - Significant biological component – hypoxia
 - Community/Political pressure to act – Executive Order (2009)

NMA – Non-Optimal Conditions

- Few data collection stations – approx. 30 online
- At each location, collect time-series of many variables (water height, velocity vector, salinity, temperature,...)
- Given a limited number of extracted modes, can NMA meet its goals in the Chesapeake Bay?
- Previous systems (Black Sea, Monterey Bay)
 - Succeeded on simpler geometries.
 - Enjoyed richer data sets (Monterey has 70% coverage).
 - Found 85% of kinetic energy populated in low numbers of modes (Black Sea).
 - Nowcasts provide current information of estuary (full domain prediction).

NMA

- Problem Statement: Can information be taken at a few select locations (under-sampled) over long periods of time (full time series) such that the spatio-temporal data set is full enough to extract amplitudes well.

Eigenmodes – features

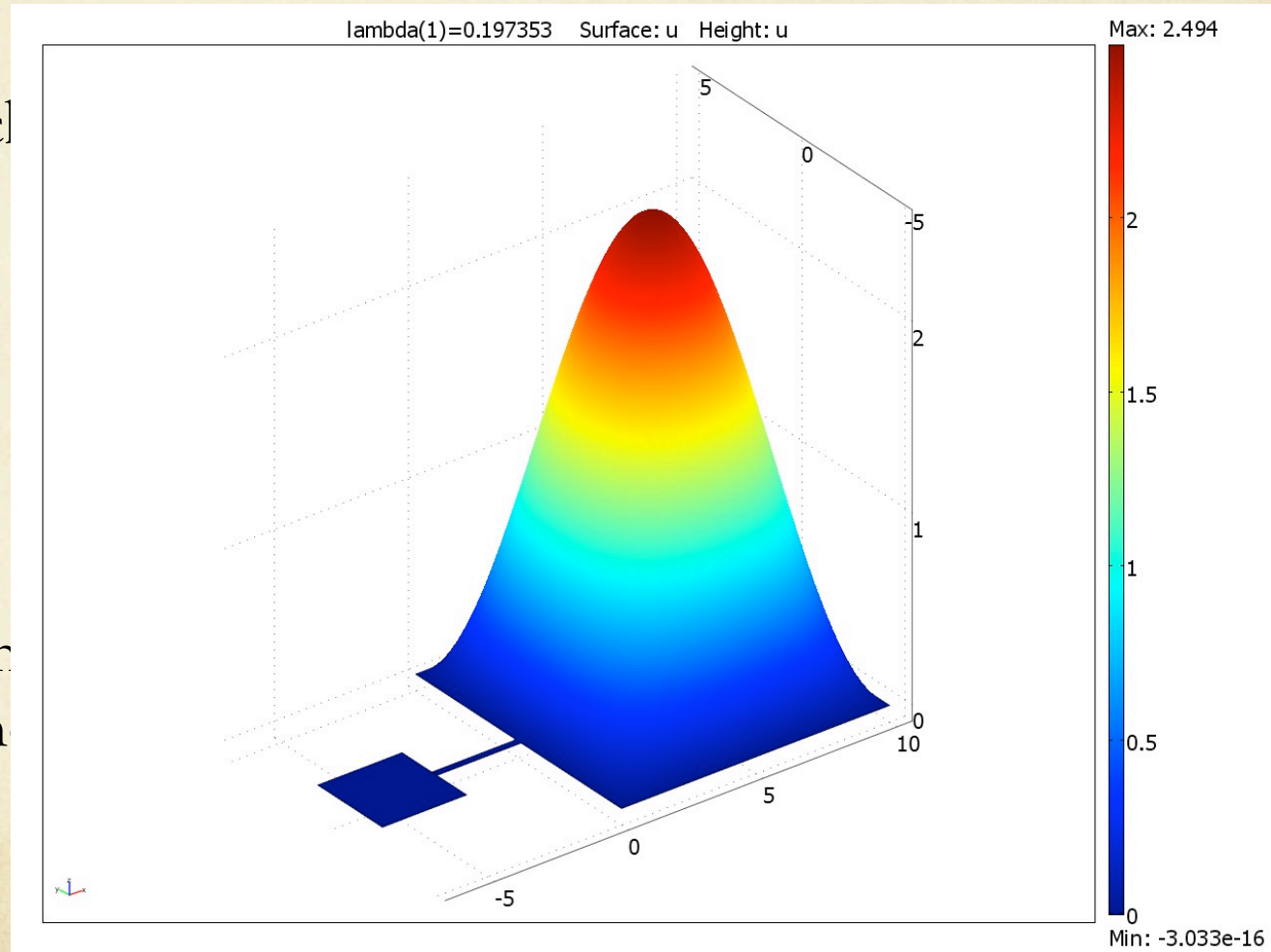
- Helmholtz Decomposition
 - Eigenmodes are the solution to Helmholtz equation.
 - Dirichlet (vorticity) / Neuman (divergence) boundary conditions used
 - Velocity vector fields calculated – $\mathbf{u}_{D,n}(\mathbf{x},t)$, $\mathbf{u}_{N,n}(\mathbf{x},t)$
 - Eigenvalues track with geometric length scales.
 - Leads to global modes and local modes.
- In order to provide 10+ eigenmodes over all regions of the Chesapeake Bay, 100+ eigenmodes need to be used.

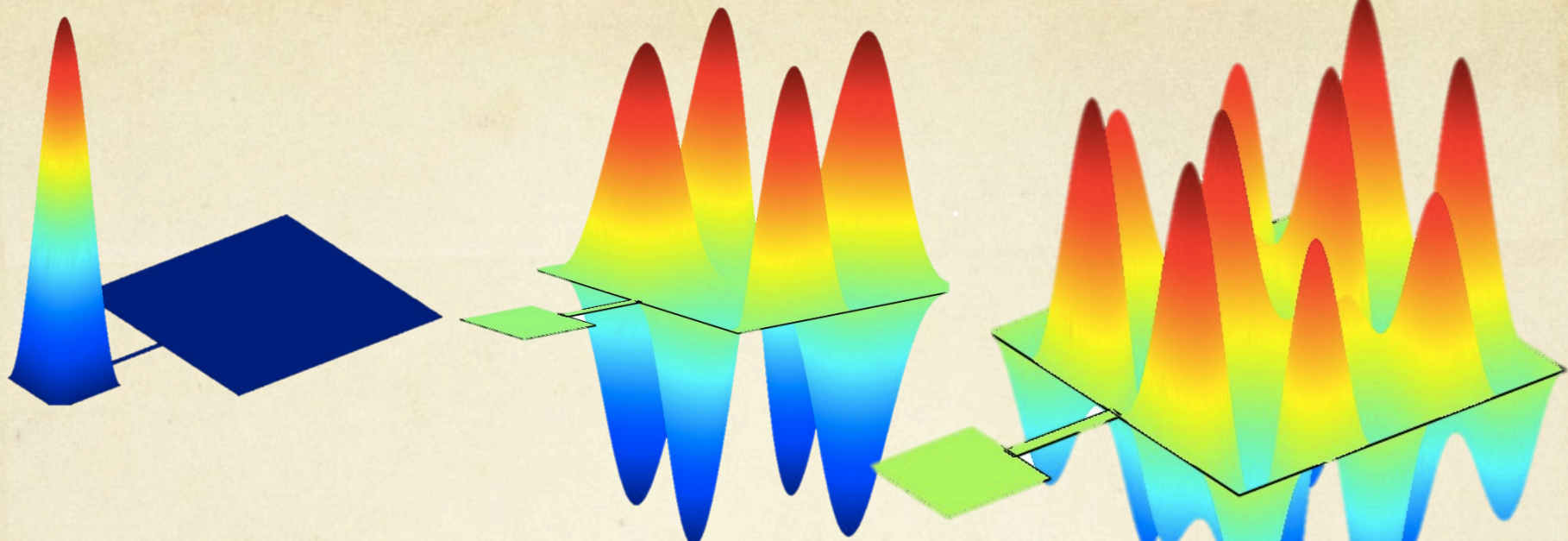
Eigenmodes - Interpretation

- A simple ortho-normal basis set (an alphabet).
- A physical mode related to dynamics of the system.
- Physical modes should exhibit differing time-scales (diffusion).
- Question: What are the appropriate time-scales associated with modal structure?
- How quickly can energy be transferred from one mode to (FPU):
 - Another mode?
 - Many modes?
 - All modes (thermalize)?

Eigenmodes – global/local

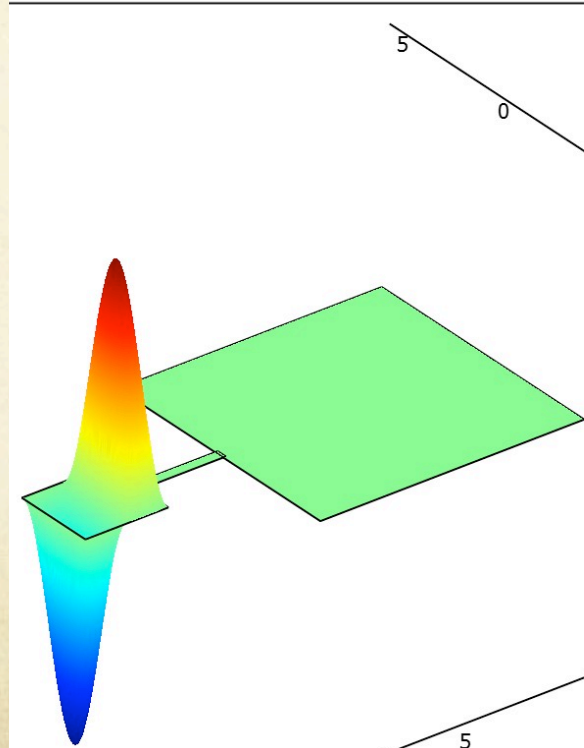
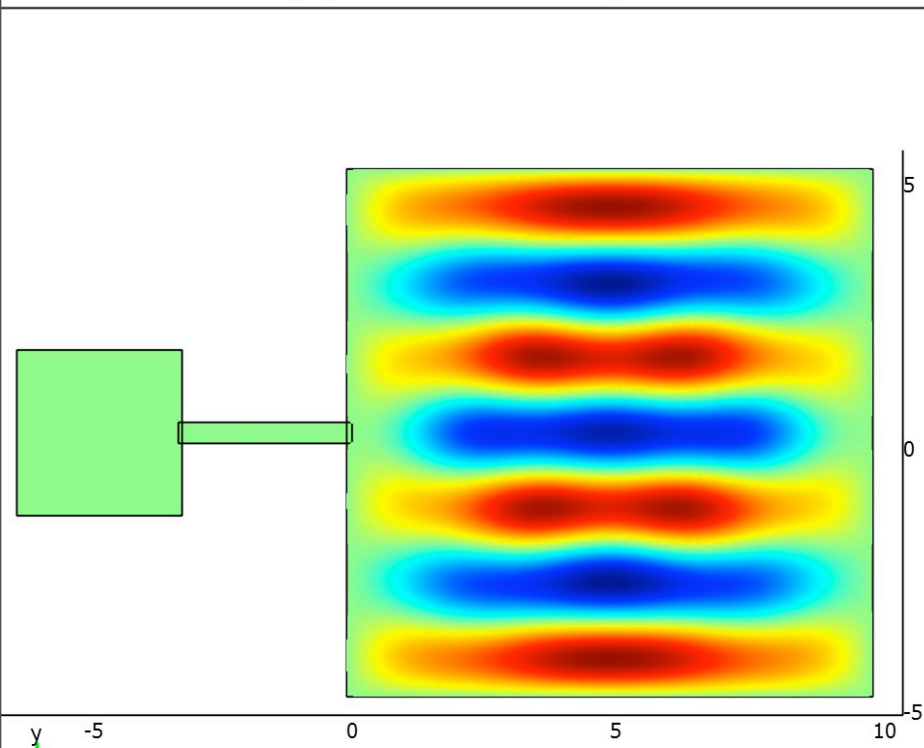
- Global modes – normal kind which cover the domain
- Local modes – only fit spatially in one location in the domain





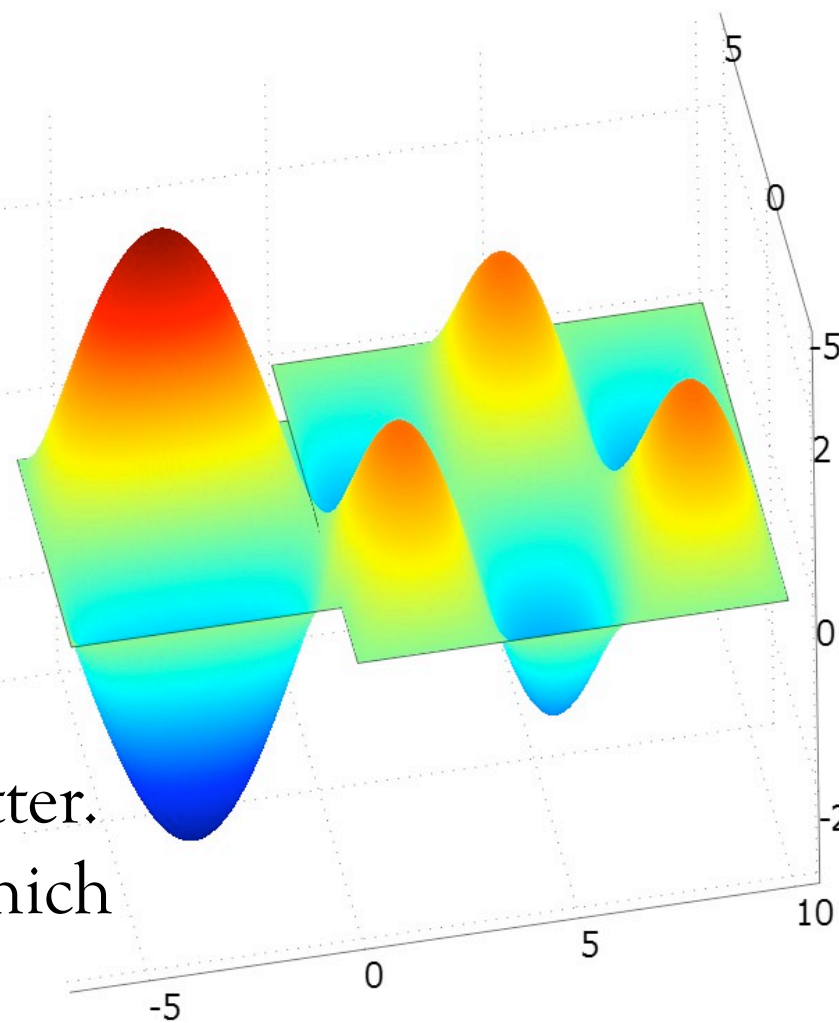
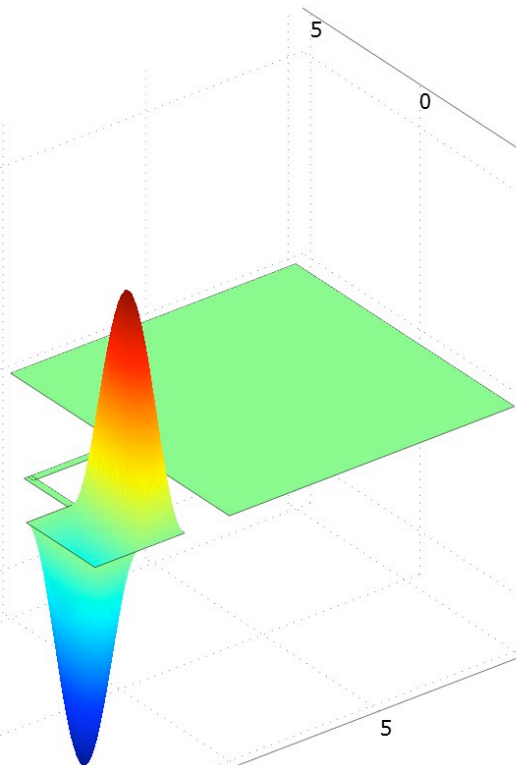
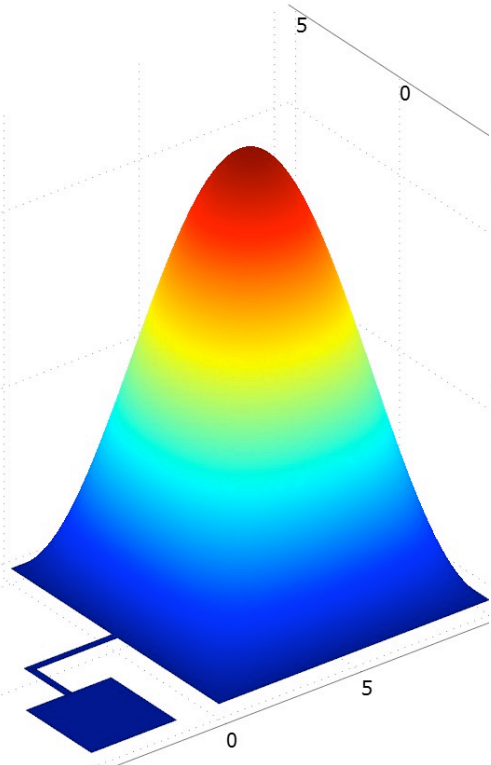
lambda(34)=4.934807 Surface: u Height: u

lambda(35)=4.956725 Surface: u Height: u

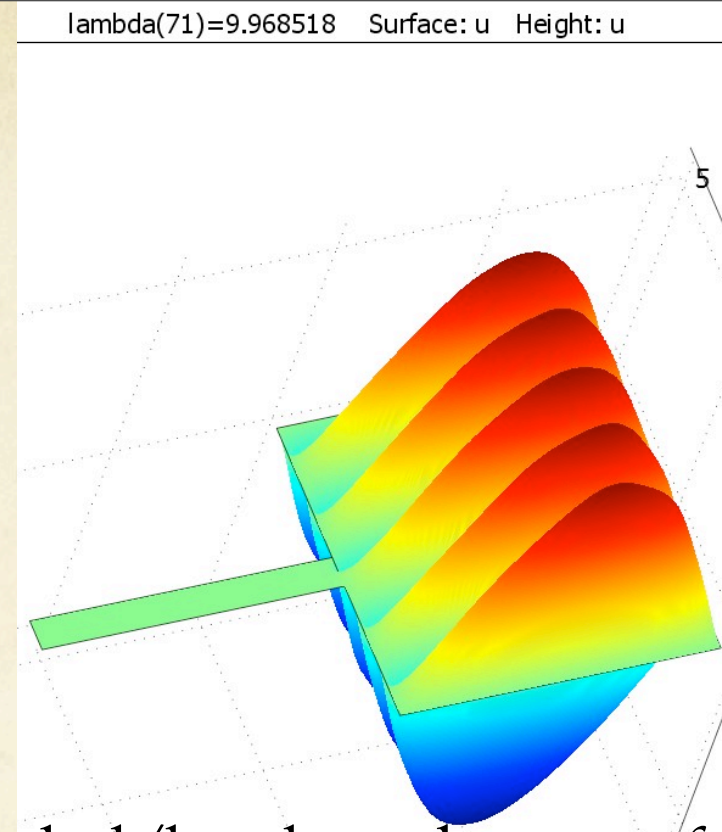
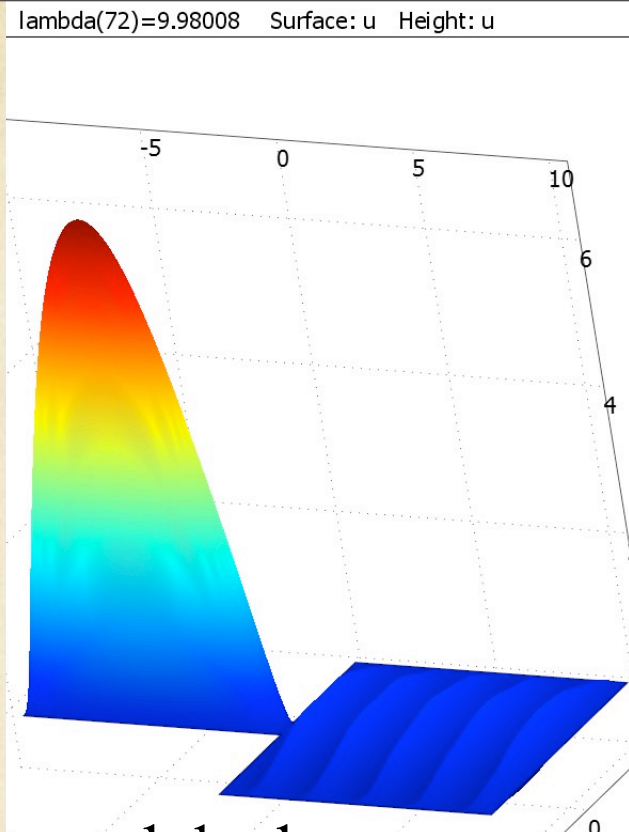


Very sensitive to eigenvalues (lengths).

$$\lambda = \pi^2 / L^2 * (n_x^2 + n_y^2)$$

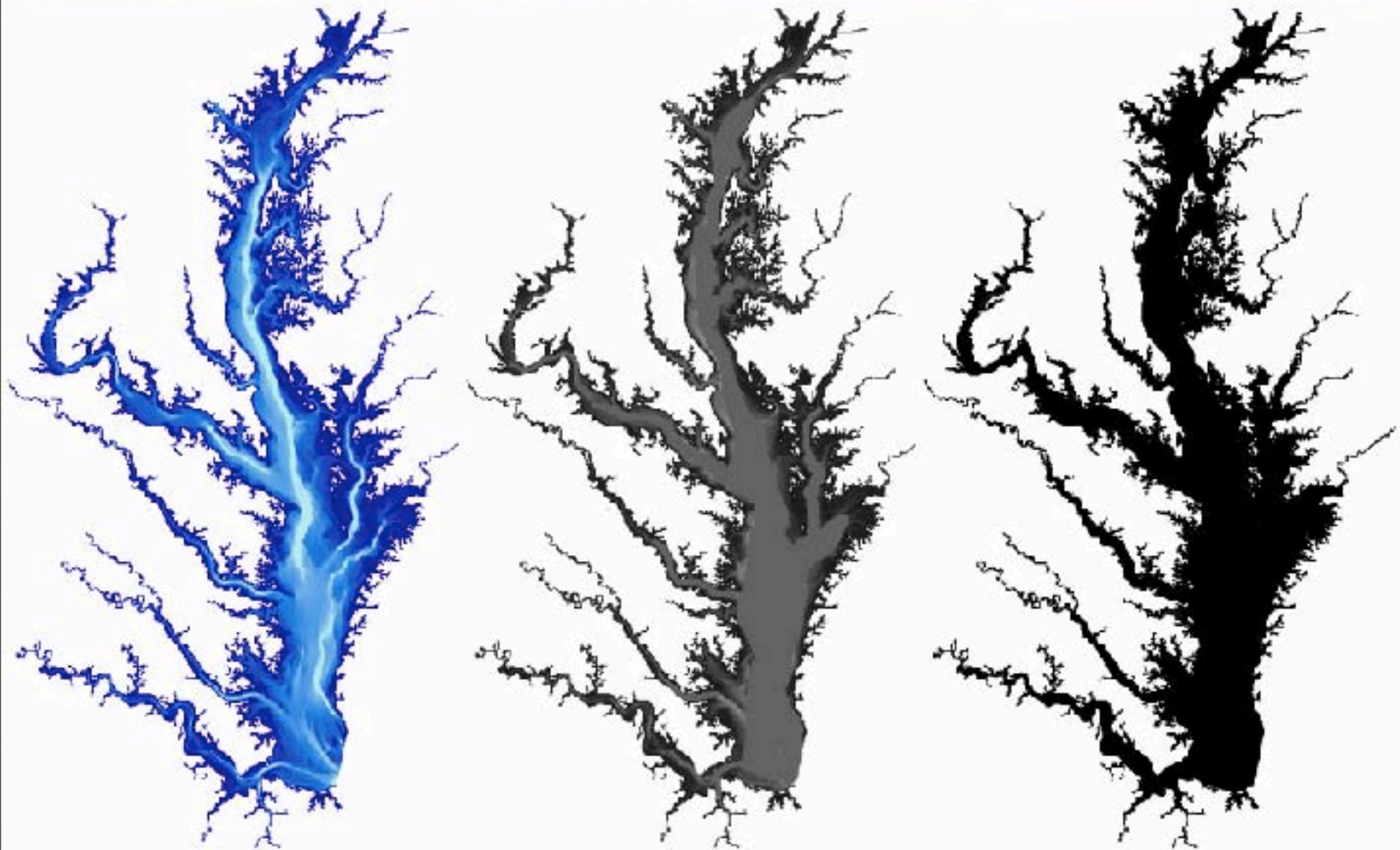


- The connecting path does not matter.
- There can be degenerate modes which are global/local.
- When allowed, modes can exist which stretch one wavelength and compress another to accommodate a compromised mode.



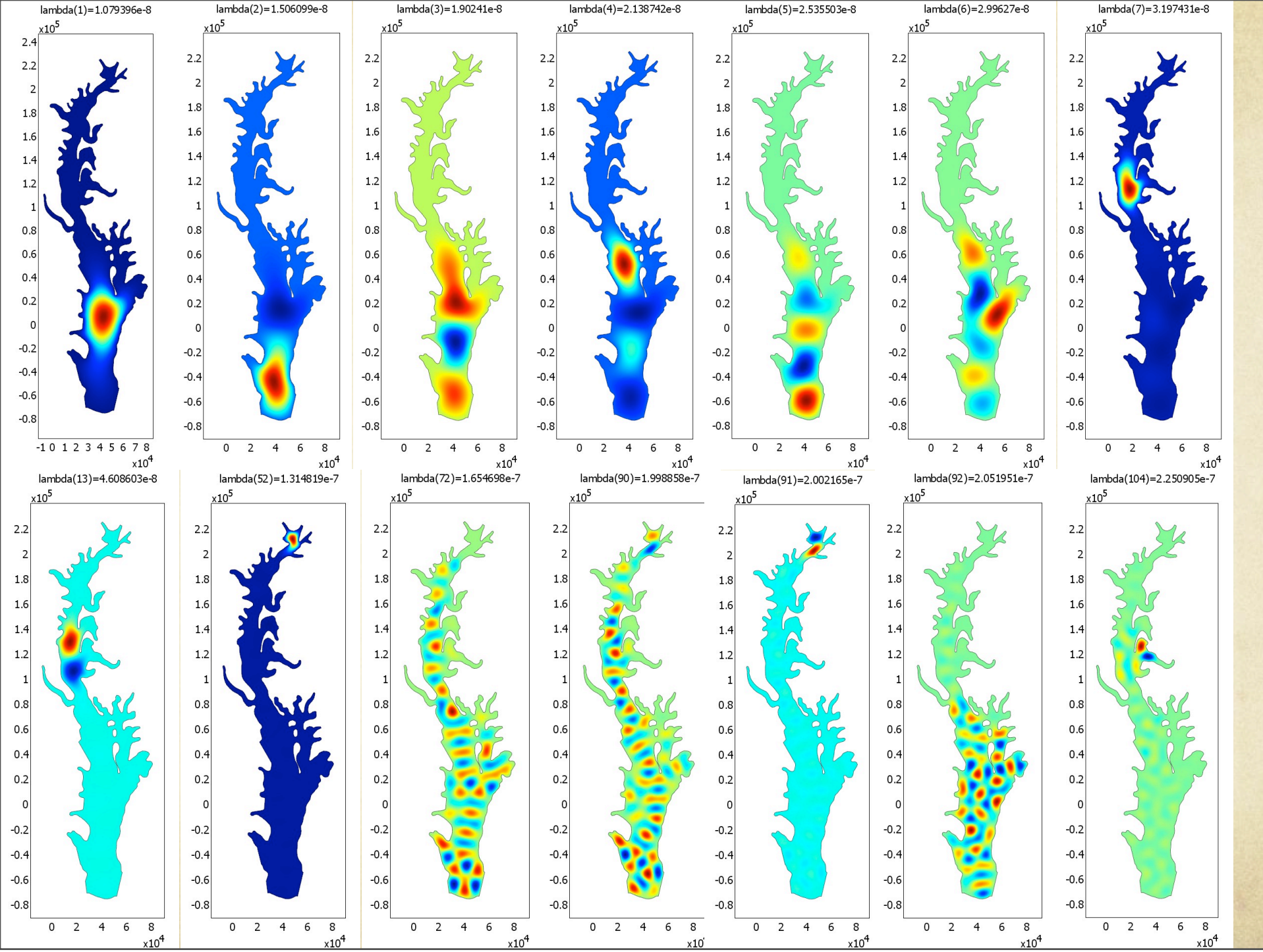
- Toy models demonstrate the global/local modes exist for many simple geometries. Such as:
 - Two geometries “connected” by waterway.
 - Bay/Ocean.
 - River/Bay.
- Global/local mode structures should be expected whenever lengths cover multi-scales.

Image Processing of the Chesapeake

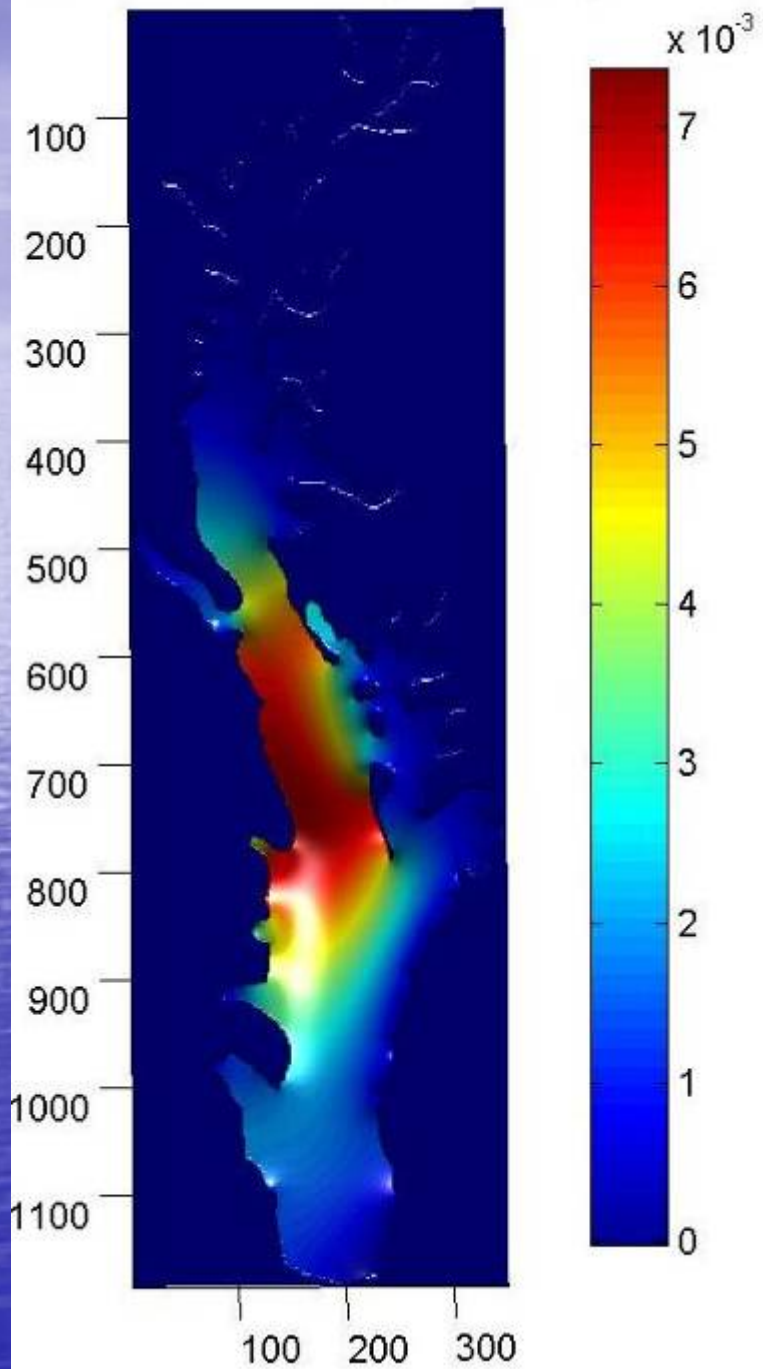


Approximated Boundaries

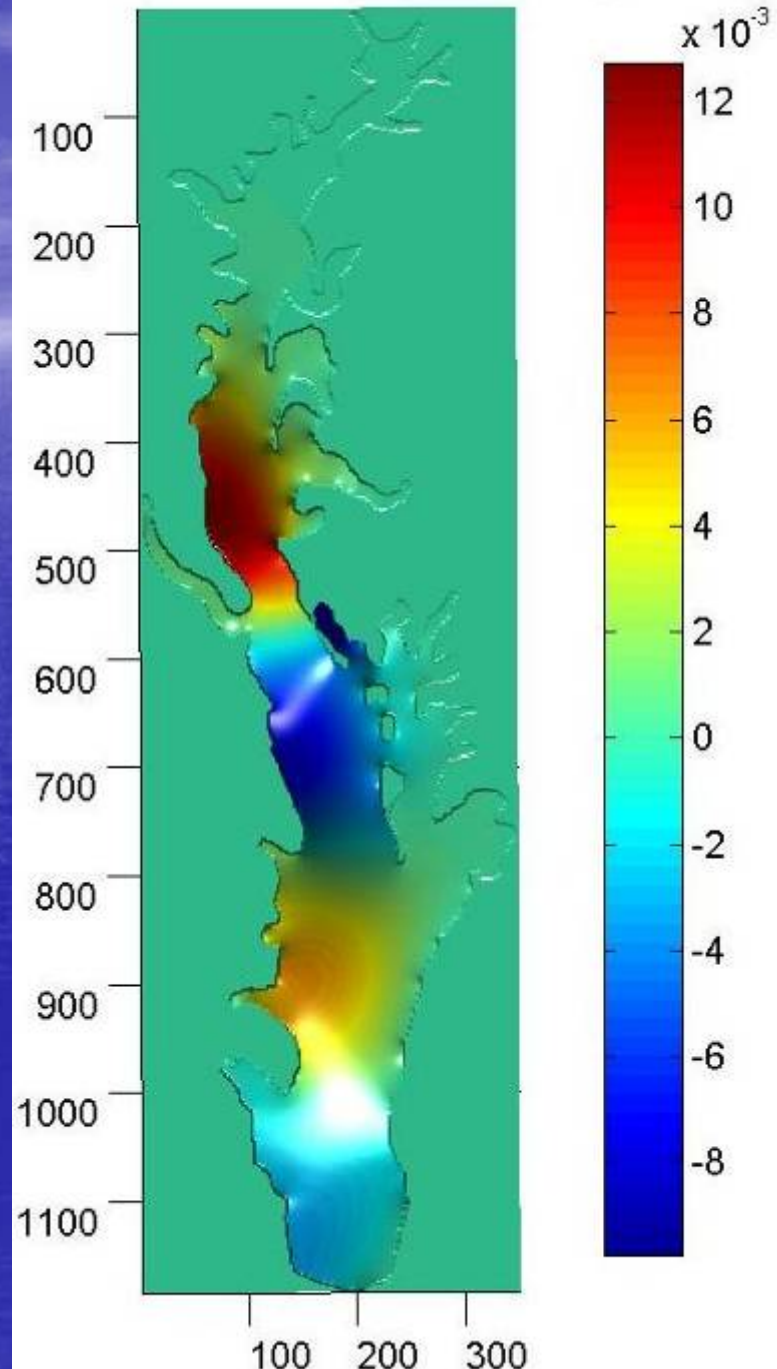




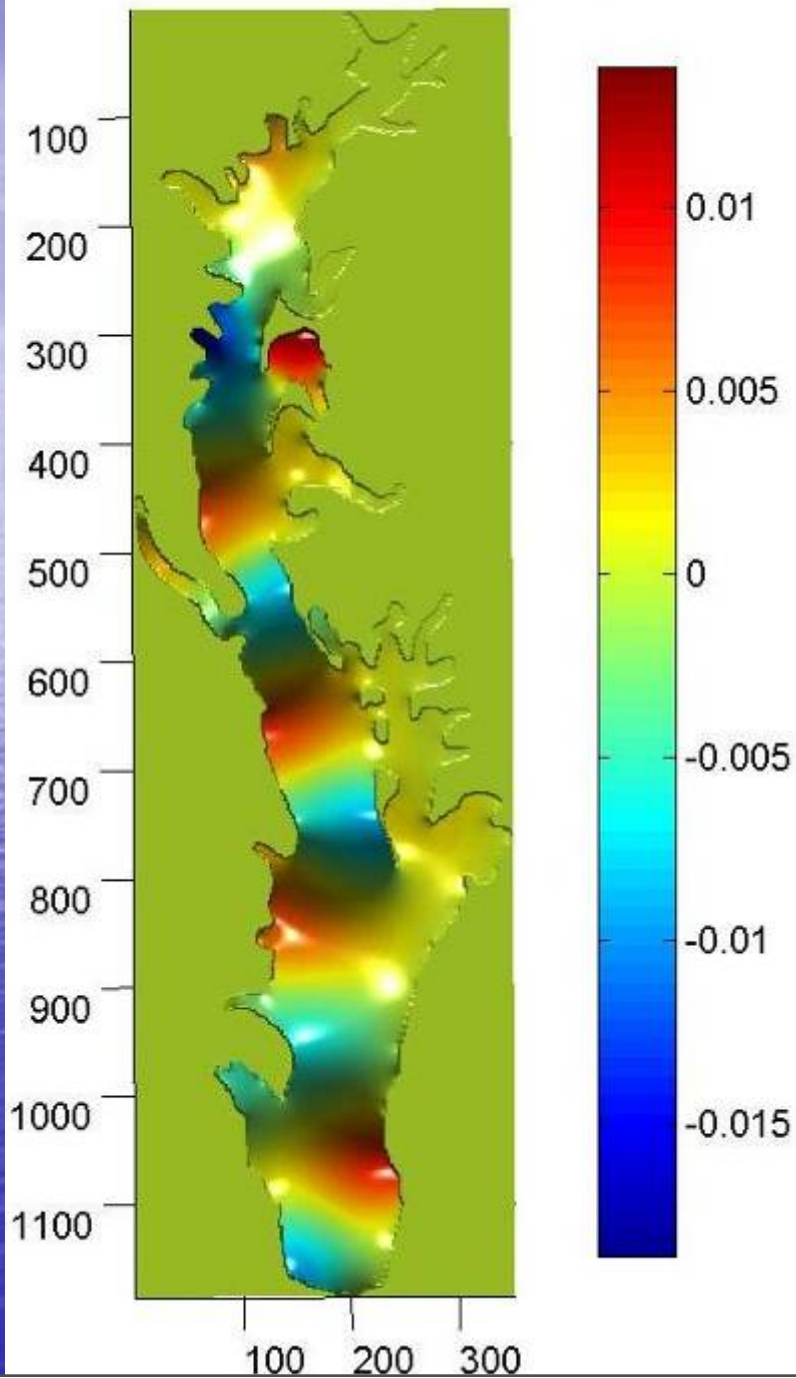
Neumann Mode 1 (350X1185)



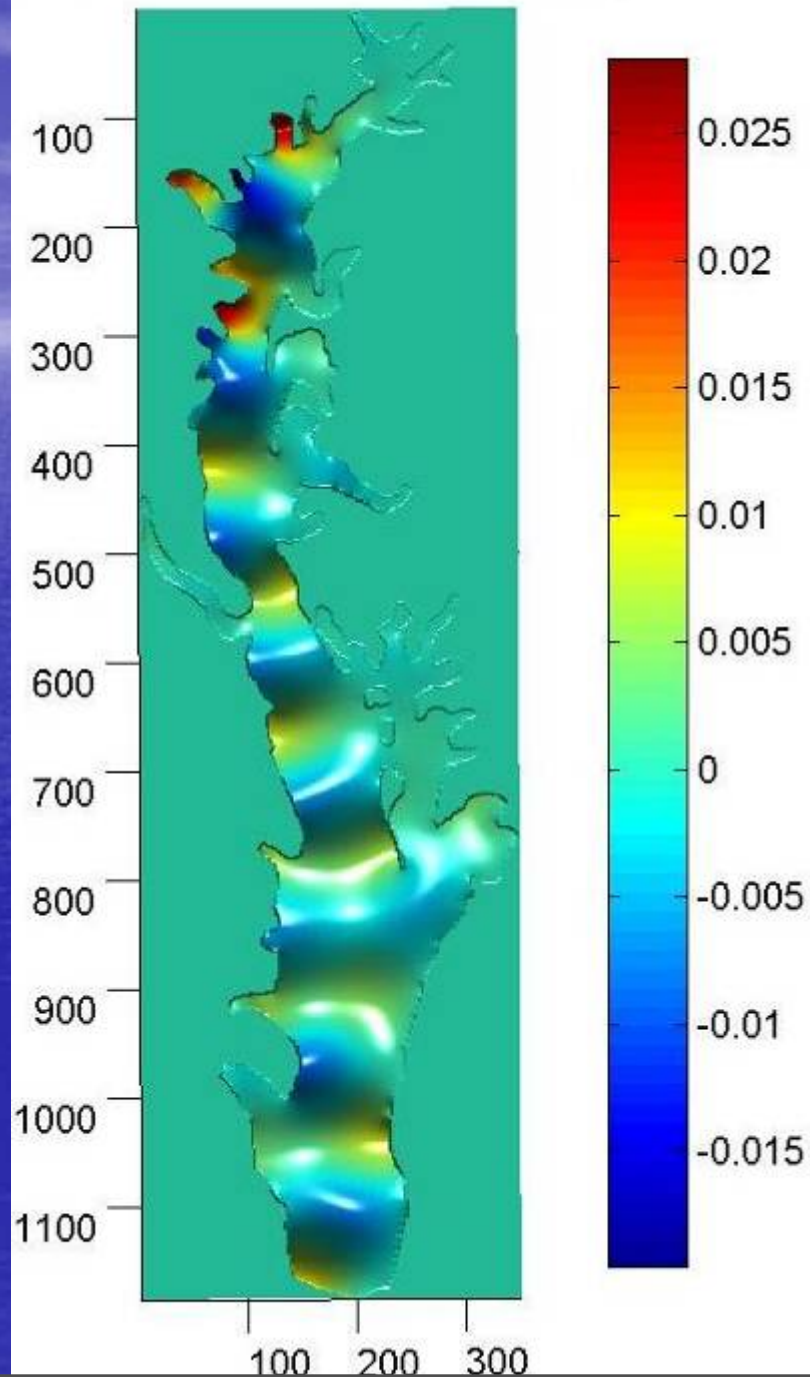
Neumann Mode 4 (350X1185)



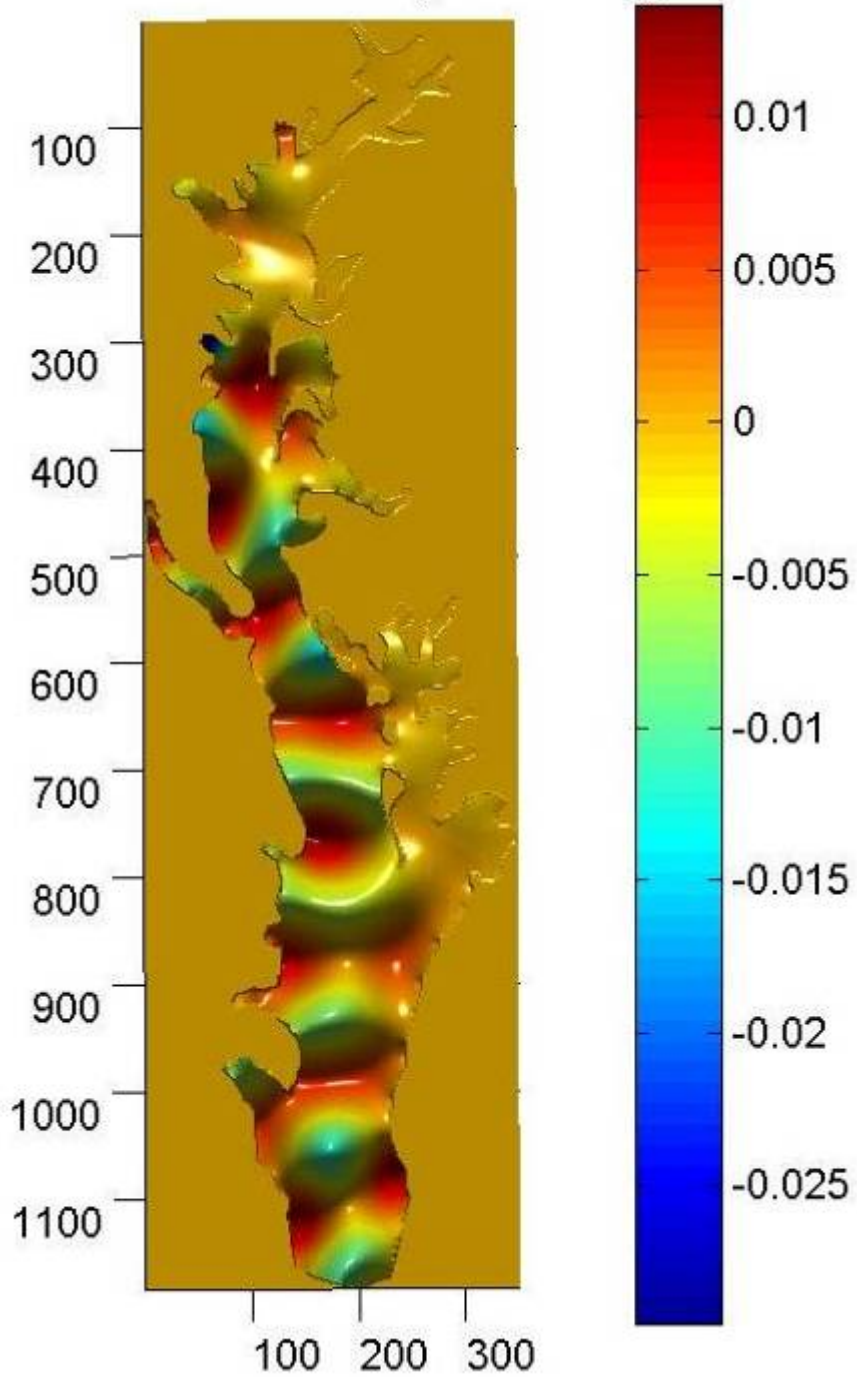
Neumann Mode 15 (350X1185)



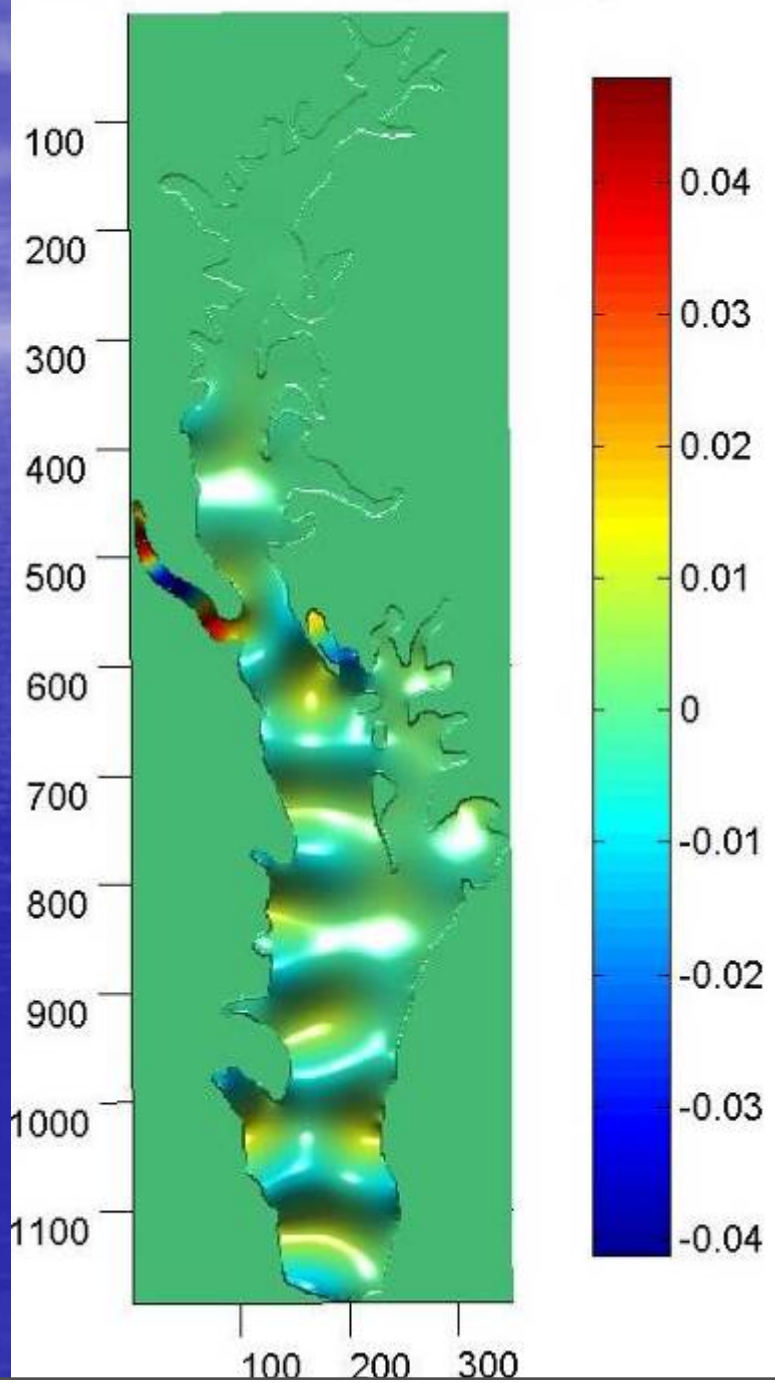
Neumann Mode 34 (350X1185)



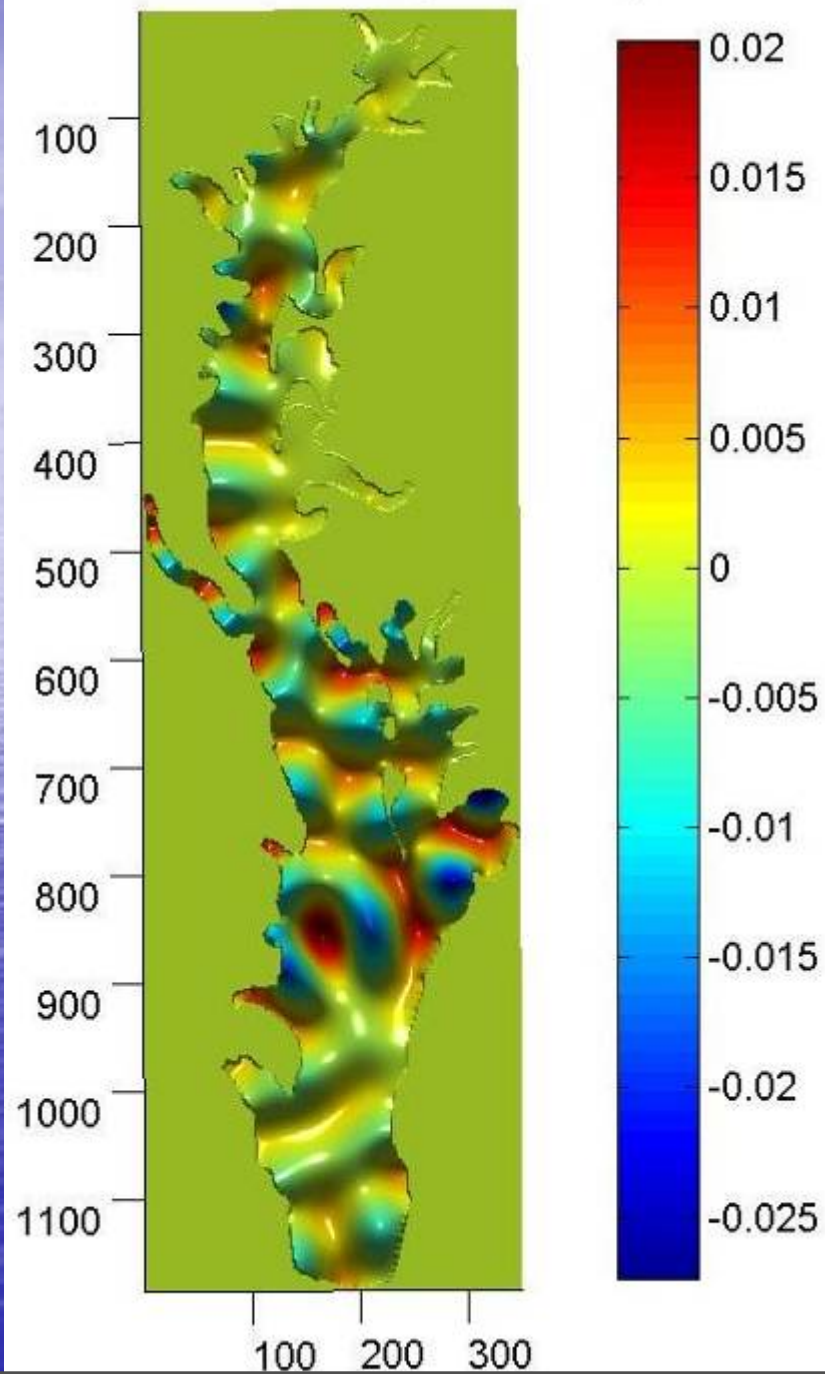
Neumann Mode 48 (350X1185)



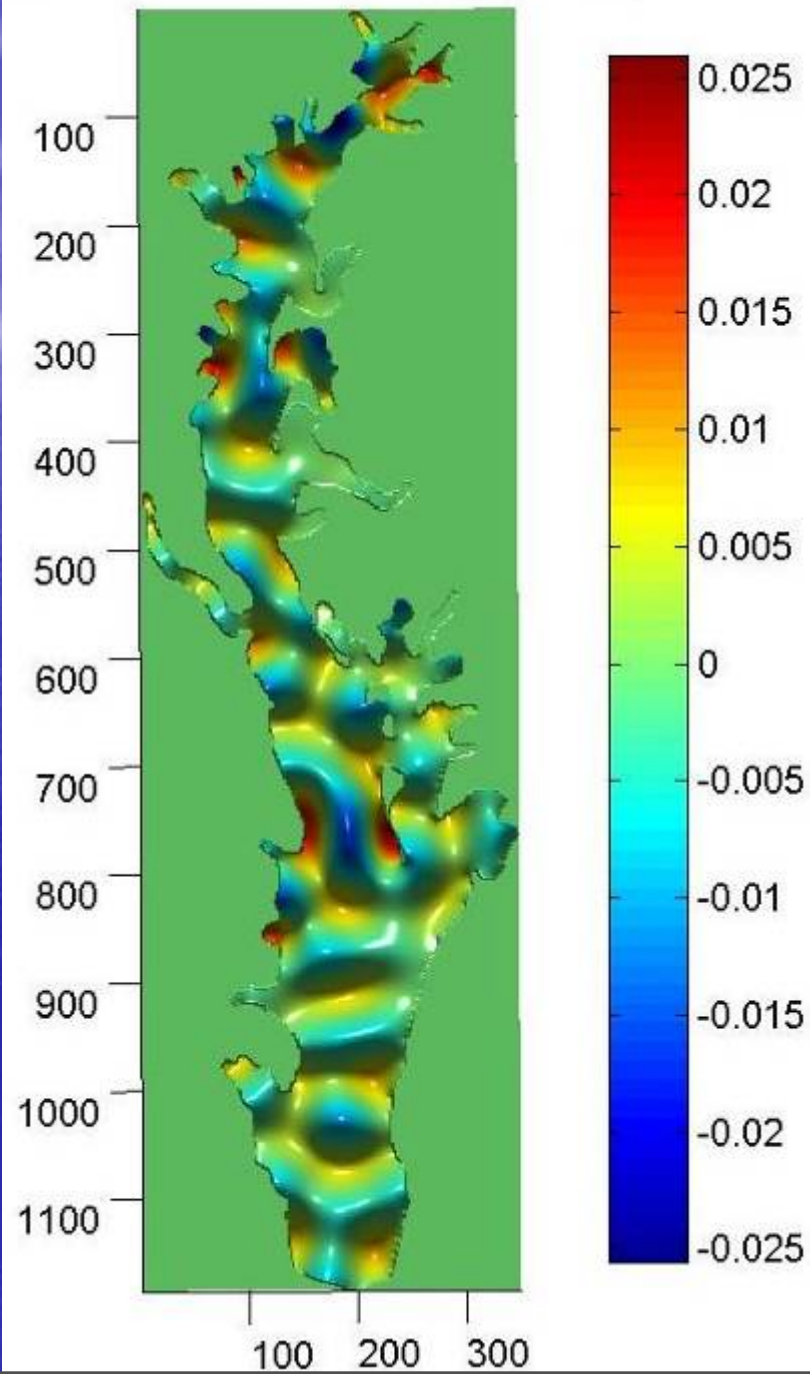
Neumann Mode 58 (350X1185)



Neumann Mode 91 (350X1185)



Neumann Mode 100 (350X1185)

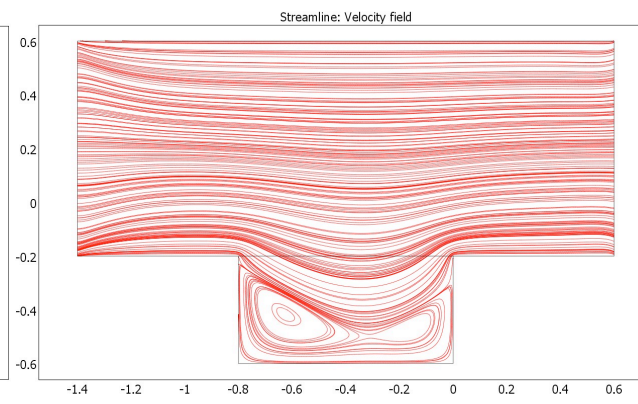
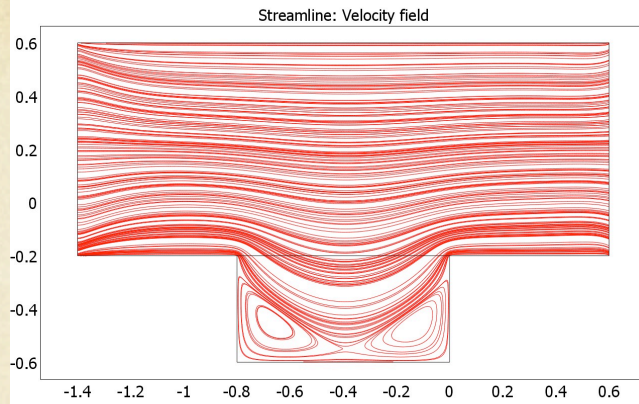


Cavity Driven Flow

- By increasing the forcing along the main channel, do any higher modes ever occur other than lowest?
- How fast can one mode “relax” into other modes?
- In order for physical gyres to be found, DHTs are required to *guide* the BVP.

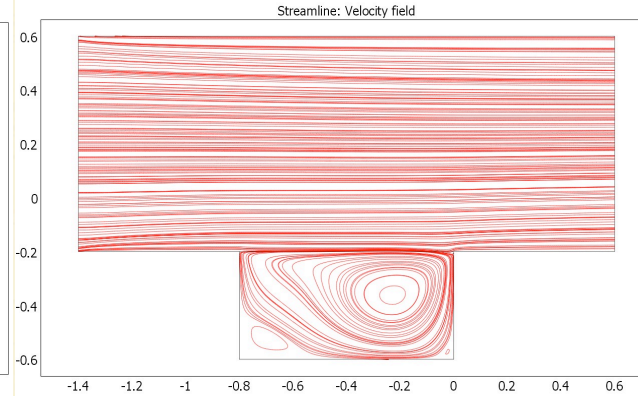
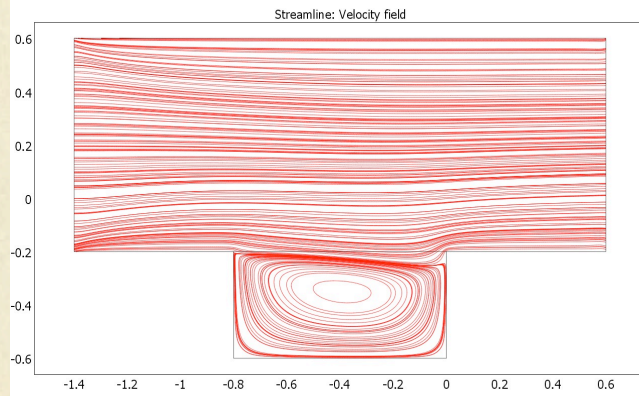
INLET:

$V=1$ m/s



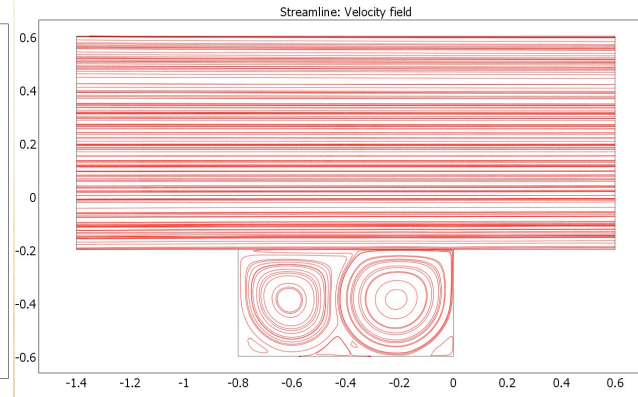
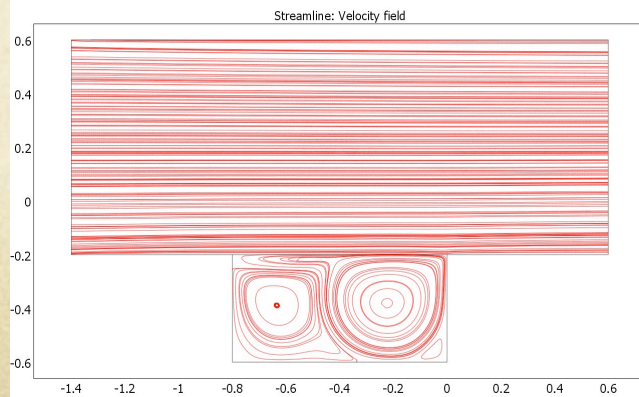
$V = 10$ m/s

$V= 100$ m/s



$V = 1000$ m/s

$V=10000$ m/s

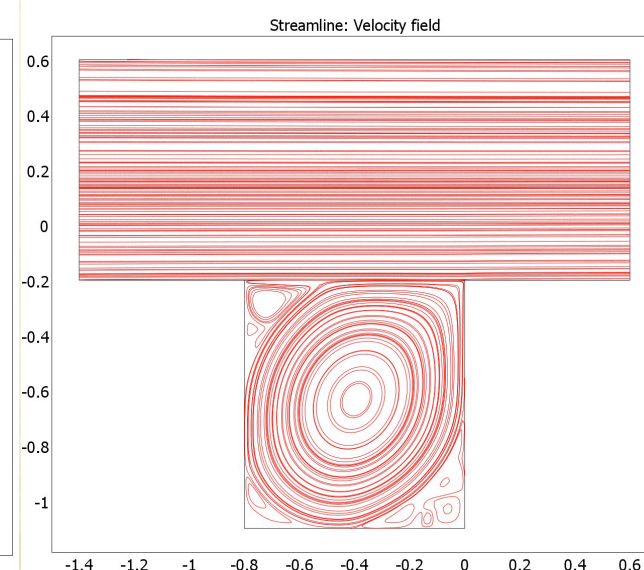
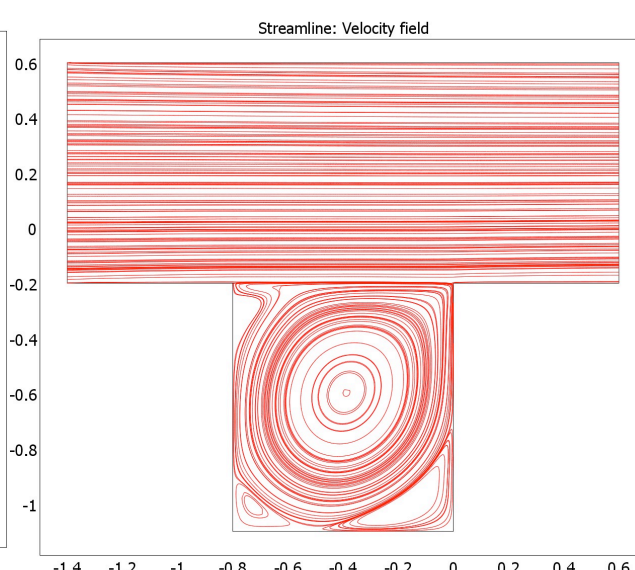
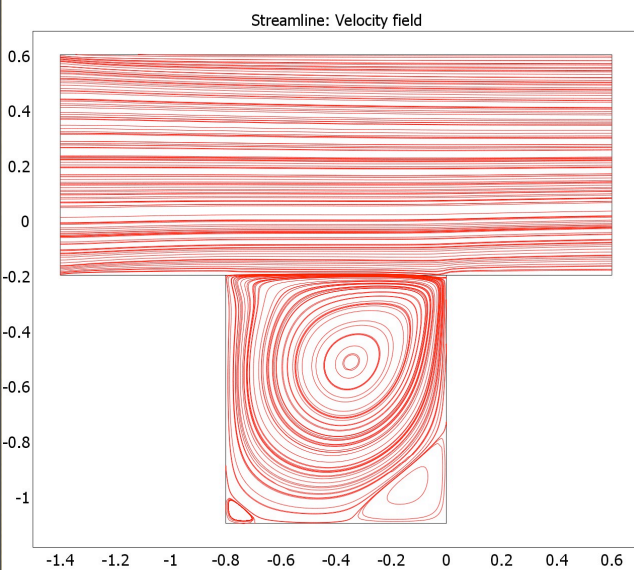
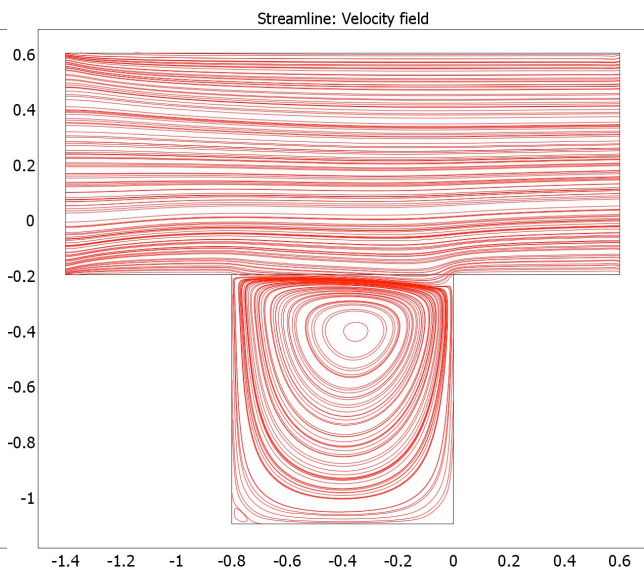
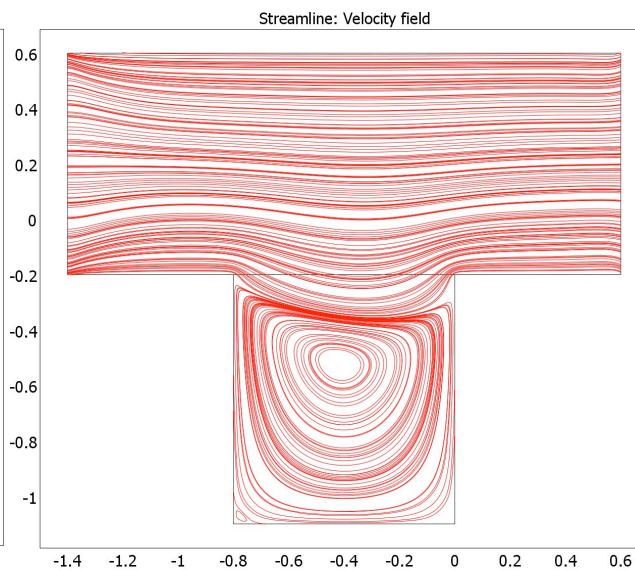
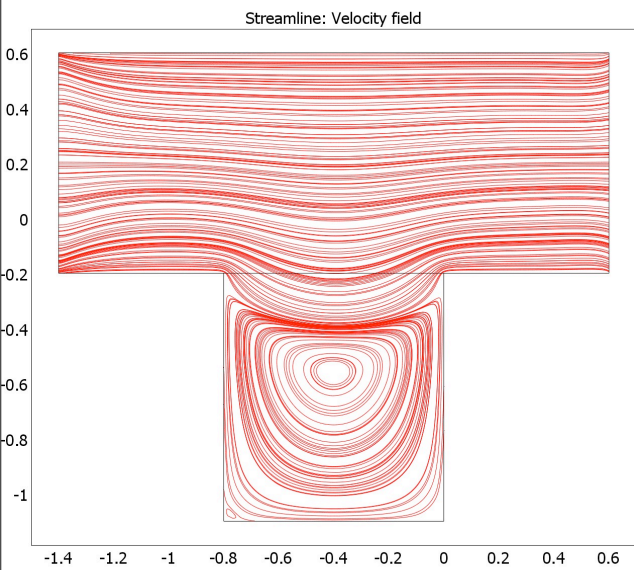


$V=10^6$ m/s

INLET: $V=1$ m/s

$V=10$ m/s

$V=100$ m/s

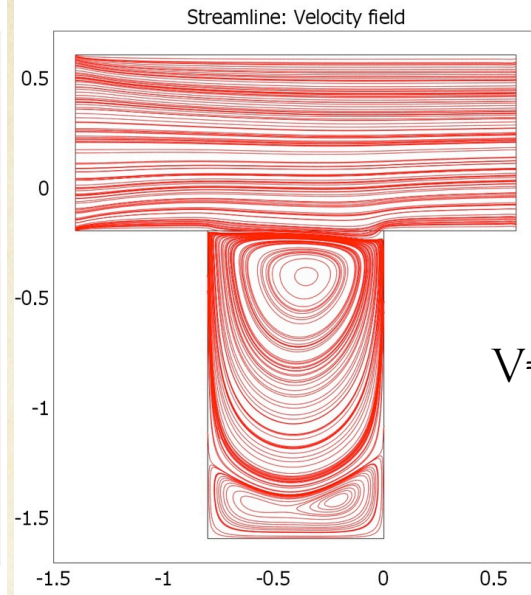
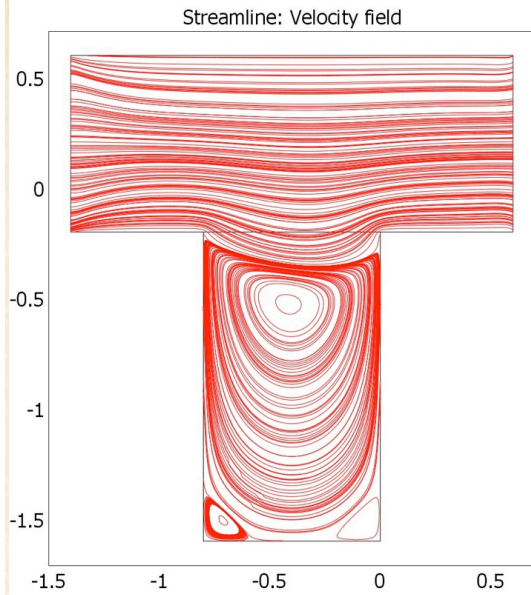
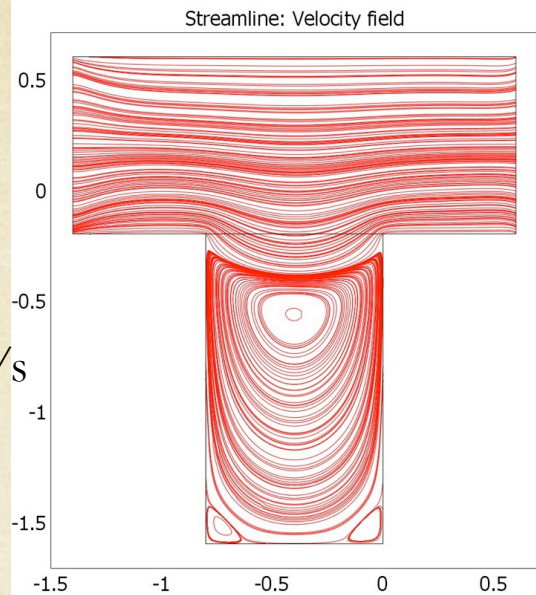


INLET: $V=1000$ m/s

$V=10000$ m/s

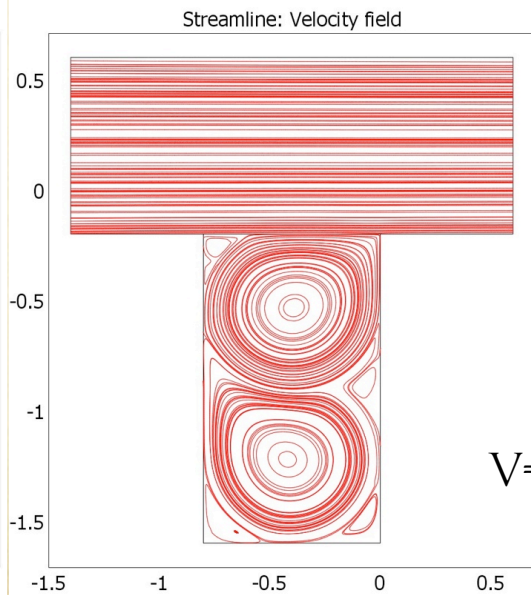
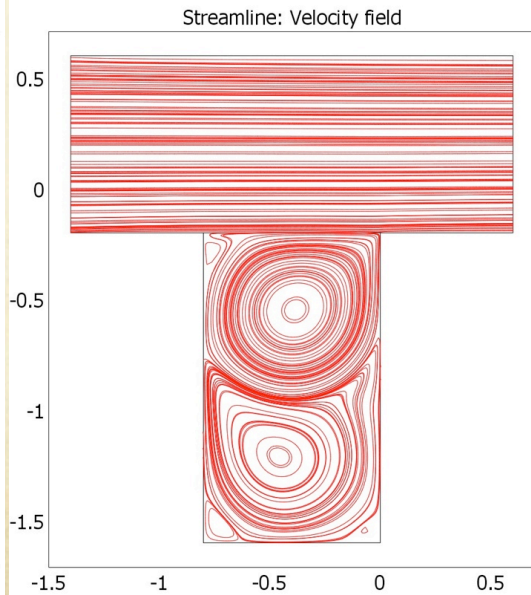
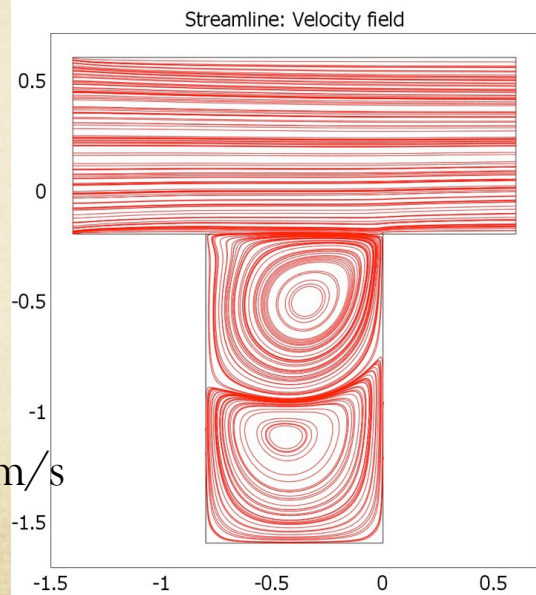
$V=10^6$ m/s

$V=1\text{m/s}$



$V=10^2\text{m/s}$

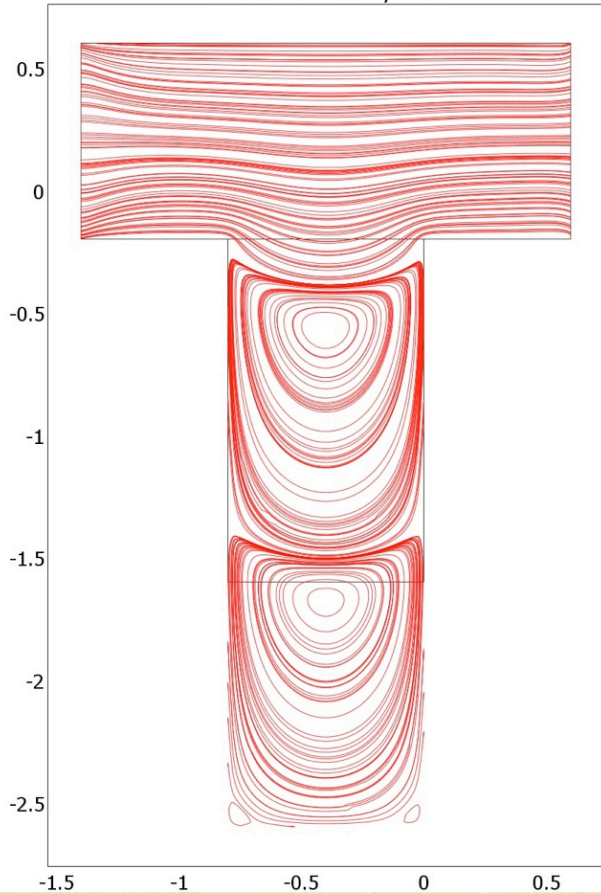
$V=10^3\text{m/s}$



$V=10^5\text{m/s}$

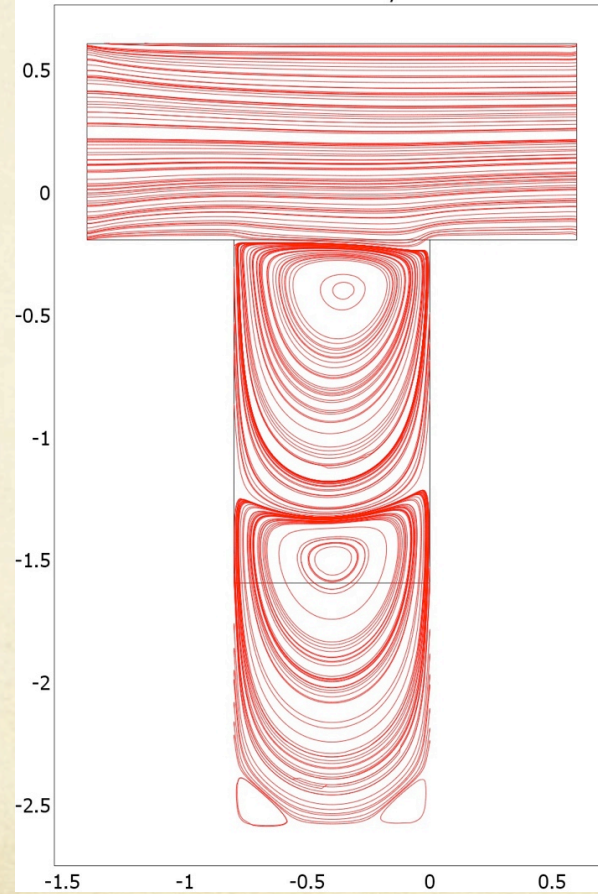
3 Modes Deep

Streamline: Velocity field



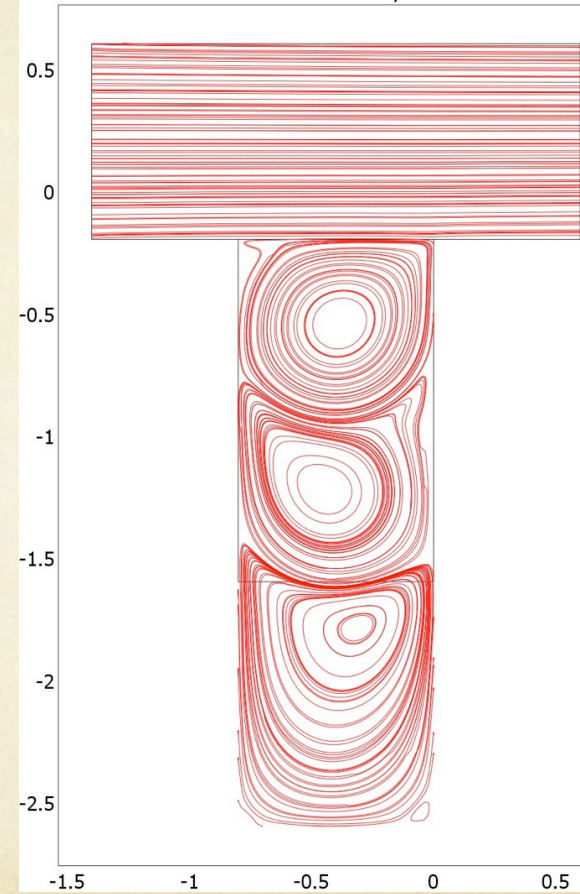
$V = 1 \text{ m/s}$

Streamline: Velocity field



$V = 100 \text{ m/s}$

Streamline: Velocity field



$V = 10000 \text{ m/s}$

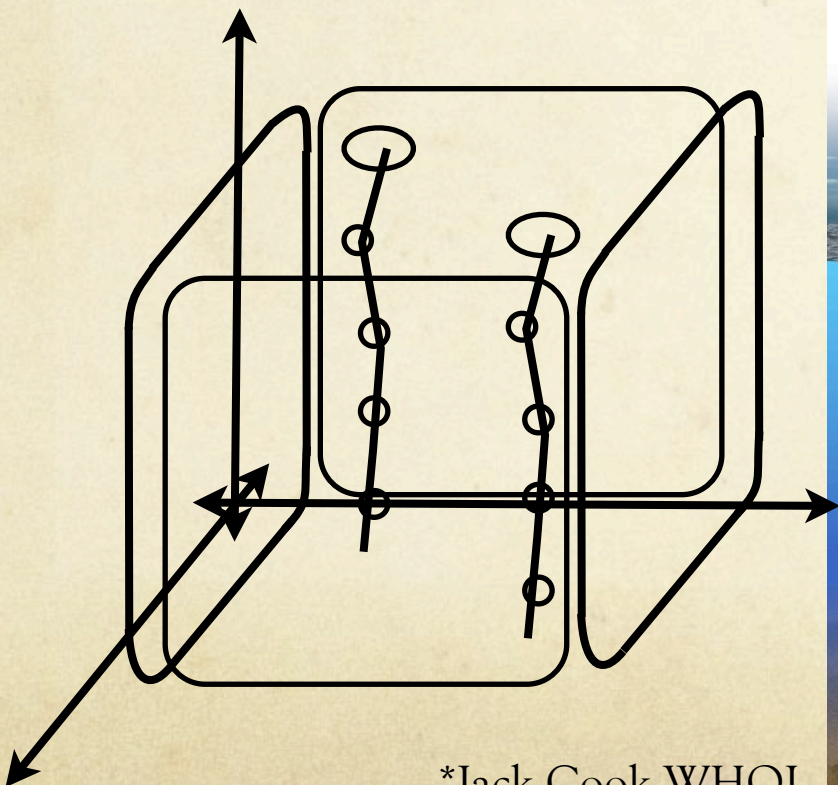
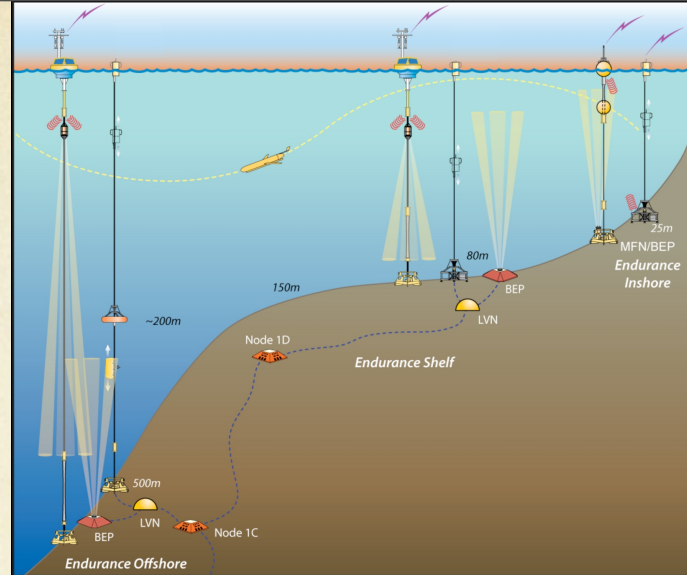
Data Collection for 3D modes

- Assume most of data is collected regularly on the surface of the water.
- At a limited set of locations, data is collected at depth down to a limited depth and for a finite number of depths.
- Lagrangian drifter/gliders data exists along transects of the domain (specific to one location at one time).
- Satellite feature extraction only works when coverage permits.

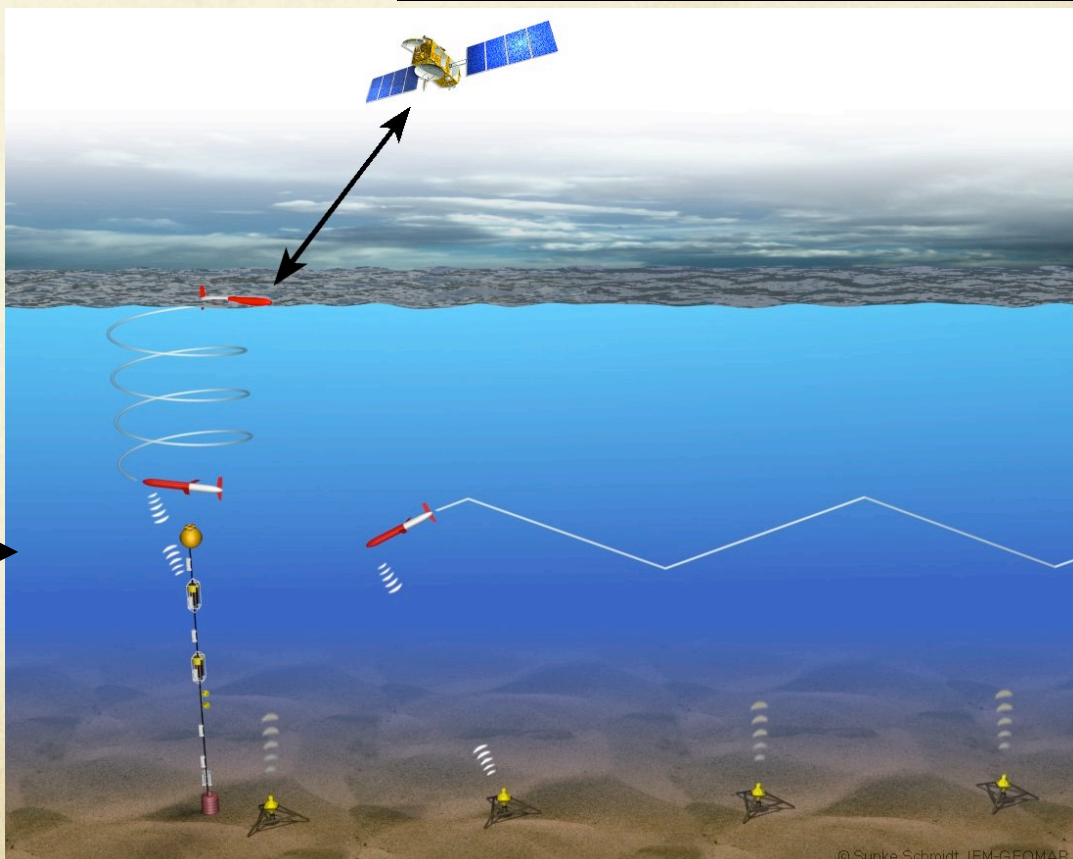
Most of this data will be taken on the top layer (surface).

Some data at depth is taken.

Glider data taken along a transect.



*Jack Cook WHOI

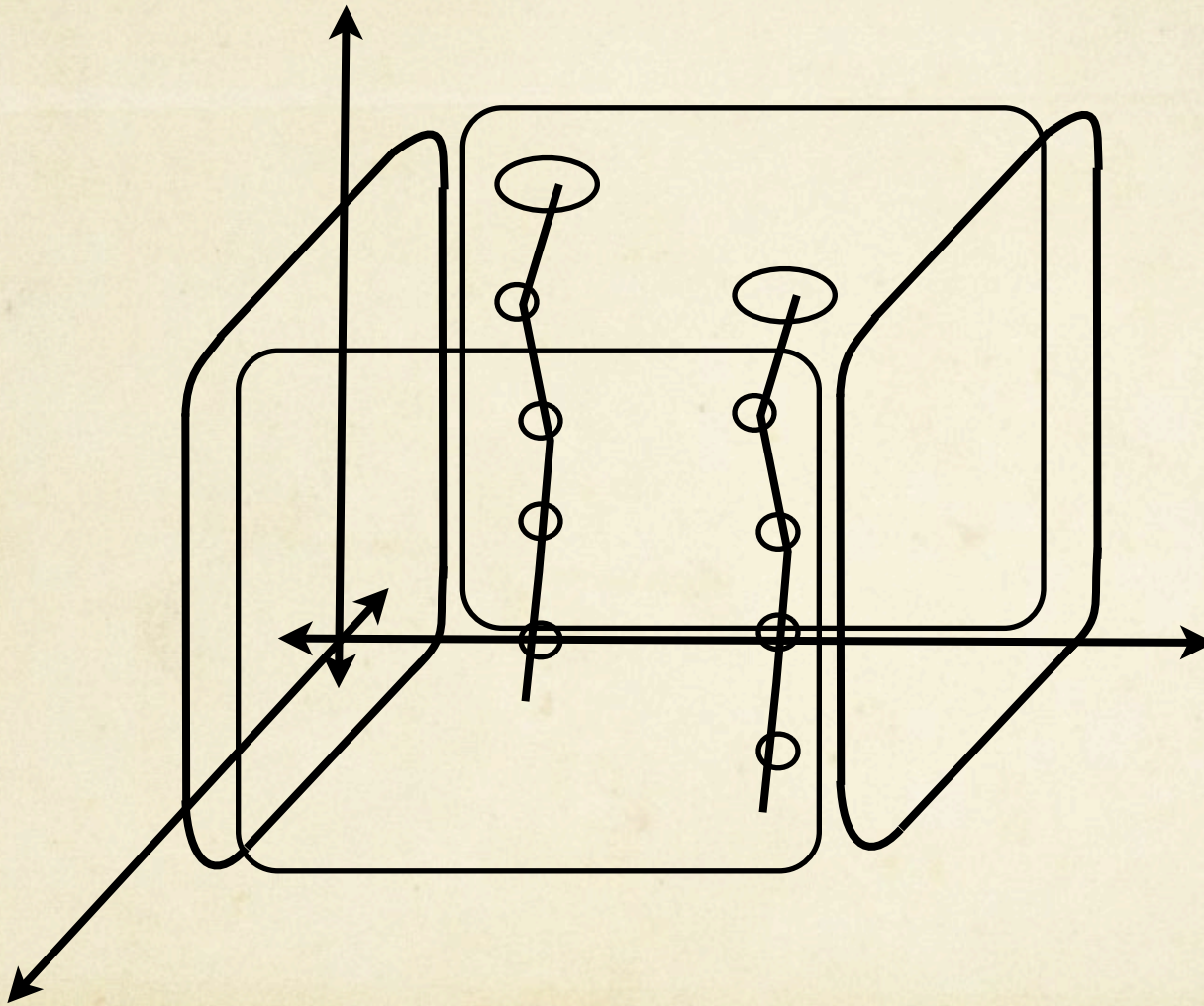


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3D Modes - limitations

- Most of the data is collected on one face of a 3D mode.
- Coverage of the face will be limited (less than 100%).
- A few lines along at depth (less than 10% coverage).
- Drifter/gliders data force issue of how to incorporate spatio-temporal information in a meaningful way to extract amplitudes of modes.
- Shannon sampling theorem – need as many spatial points as modes to extract (more or less).
- Compressive sensing addresses this partially.
- Can times-series data be used as additional data points – effectively increasing the size of the data set spatially – providing enough coverage to be useful?

3D Mode Data Collection



Most of this data will be taken on the top layer (surface).

Some data at depth is taken.

Possible Solution

- Collect data at a limited number of locations/times

$$f(\mathbf{r}_i, t_j) \quad \mathbf{u}(\mathbf{r}_i, t_j)$$

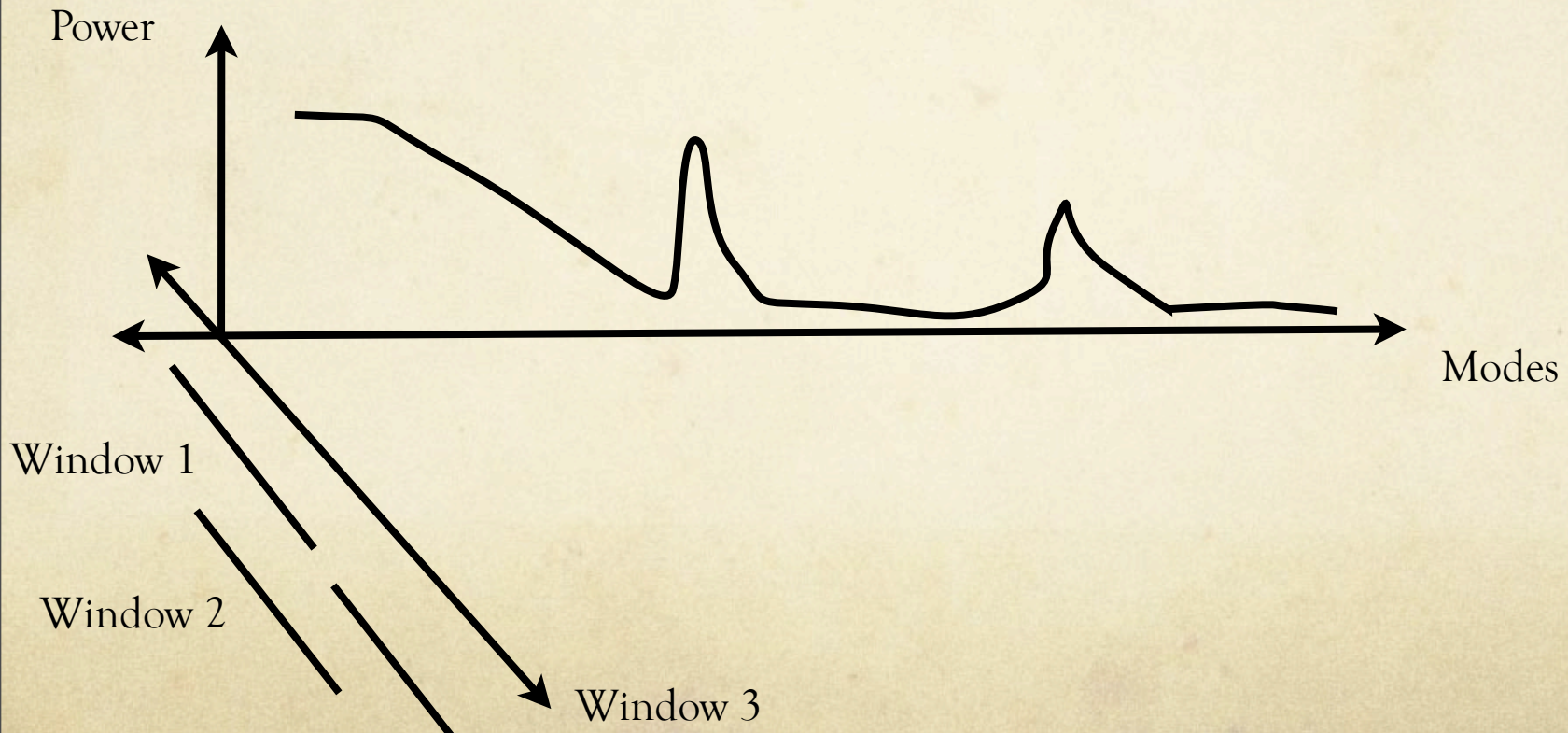
- Select a set of times for a window (N_t in window)

$$t_1 \dots t_j \dots t_{N_t}$$

- Treat all data within the time window as if constant wrt. amplitudes of the modes.
- Fit the amplitudes (A_n, B_n) to the extended data set.

Power Spectrum over Time

- Assume spectrum is constant over a time window
- Calculate the spectra over multiple overlapping windows.



Fitting the Data

$$\begin{pmatrix} u(\mathbf{r}_1, t_1) \\ u(\mathbf{r}_1, t_2) \\ u(\mathbf{r}_1, t_3) \\ \dots \\ u(\mathbf{r}_1, t_{N_t}) \\ v(\mathbf{r}_1, t_1) \\ v(\mathbf{r}_1, t_2) \\ v(\mathbf{r}_1, t_3) \\ \dots \\ v(\mathbf{r}_1, t_{N_t}) \\ u(\mathbf{r}_2, t_1) \\ \dots \\ u(\mathbf{r}_2, t_{N_t}) \\ v(\mathbf{r}_2, t_1) \\ \dots \\ v(\mathbf{r}_2, t_{N_t}) \\ \dots \\ u(\mathbf{r}_{N_p}, t_1) \\ \dots \\ u(\mathbf{r}_{N_p}, t_{N_t}) \\ v(\mathbf{r}_{N_p}, t_1) \\ \dots \\ v(\mathbf{r}_{N_p}, t_{N_t}) \end{pmatrix} = \begin{pmatrix} \frac{\partial}{\partial y} \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) \dots \psi_N(\mathbf{r}_1) & \frac{\partial}{\partial x} \phi_1(\mathbf{r}_1) & \phi_2(\mathbf{r}_1) \dots \phi_N(\mathbf{r}_1) \\ \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) \dots \psi_N(\mathbf{r}_1) & \phi_1(\mathbf{r}_1) & \phi_2(\mathbf{r}_1) \dots \phi_N(\mathbf{r}_1) \\ \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) \dots \psi_N(\mathbf{r}_1) & \phi_1(\mathbf{r}_1) & \phi_2(\mathbf{r}_1) \dots \phi_N(\mathbf{r}_1) \\ \dots & \dots & \dots & \dots \\ \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) \dots \psi_N(\mathbf{r}_1) & \phi_1(\mathbf{r}_1) & \phi_2(\mathbf{r}_1) \dots \phi_N(\mathbf{r}_1) \\ -\frac{\partial}{\partial x} \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) \dots \psi_N(\mathbf{r}_1) & \frac{\partial}{\partial y} \phi_1(\mathbf{r}_1) & \phi_2(\mathbf{r}_1) \dots \phi_N(\mathbf{r}_1) \\ \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots \\ \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) \dots \psi_N(\mathbf{r}_1) & \phi_1(\mathbf{r}_1) & \phi_2(\mathbf{r}_1) \dots \phi_N(\mathbf{r}_1) \\ \psi_1(\mathbf{r}_2) & \psi_2(\mathbf{r}_2) \dots \psi_N(\mathbf{r}_2) & \phi_1(\mathbf{r}_2) & \phi_2(\mathbf{r}_2) \dots \phi_N(\mathbf{r}_2) \\ \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots \\ \psi_1(\mathbf{r}_2) & \psi_2(\mathbf{r}_2) \dots \psi_N(\mathbf{r}_2) & \phi_1(\mathbf{r}_2) & \phi_2(\mathbf{r}_2) \dots \phi_N(\mathbf{r}_2) \\ \psi_1(\mathbf{r}_{N_p}) & \psi_2(\mathbf{r}_{N_p}) \dots \psi_N(\mathbf{r}_{N_p}) & \phi_1(\mathbf{r}_{N_p}) & \phi_2(\mathbf{r}_{N_p}) \dots \phi_N(\mathbf{r}_{N_p}) \\ \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots \\ \psi_1(\mathbf{r}_{N_p}) & \psi_2(\mathbf{r}_{N_p}) \dots \psi_N(\mathbf{r}_{N_p}) & \phi_1(\mathbf{r}_{N_p}) & \phi_2(\mathbf{r}_{N_p}) \dots \phi_N(\mathbf{r}_{N_p}) \end{pmatrix} \begin{pmatrix} A_1 \\ A_2 \\ A_3 \\ \dots \\ A_N \\ B_1 \\ B_2 \\ B_3 \\ \dots \\ B_N \end{pmatrix}$$

Simplified...

$$f(\mathbf{r}, t) = \sum_{i=1..N} A_i(t) * \Psi_i(\mathbf{r})$$

$$\begin{pmatrix} f(\mathbf{r}_1, t_1) \\ f(\mathbf{r}_1, t_2) \\ f(\mathbf{r}_1, t_3) \\ \dots \\ f(\mathbf{r}_1, t_{N_t}) \\ f(\mathbf{r}_2, t_1) \\ f(\mathbf{r}_2, t_2) \\ f(\mathbf{r}_2, t_3) \\ \dots \\ f(\mathbf{r}_2, t_{N_t}) \\ \dots \\ \dots \\ f(\mathbf{r}_p, t_1) \\ f(\mathbf{r}_p, t_2) \\ f(\mathbf{r}_p, t_3) \\ \dots \\ f(\mathbf{r}_p, t_{N_t}) \end{pmatrix} = \begin{pmatrix} \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) & \dots & \psi_N(\mathbf{r}_1) \\ \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) & \dots & \psi_N(\mathbf{r}_1) \\ \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) & \dots & \psi_N(\mathbf{r}_1) \\ \dots & \dots & \dots & \dots \\ \psi_1(\mathbf{r}_1) & \psi_2(\mathbf{r}_1) & \dots & \psi_N(\mathbf{r}_1) \\ \psi_1(\mathbf{r}_2) & \psi_2(\mathbf{r}_2) & \dots & \psi_N(\mathbf{r}_2) \\ \psi_1(\mathbf{r}_2) & \psi_2(\mathbf{r}_2) & \dots & \psi_N(\mathbf{r}_2) \\ \psi_1(\mathbf{r}_2) & \psi_2(\mathbf{r}_2) & \dots & \psi_N(\mathbf{r}_2) \\ \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots \\ \dots & \dots & \dots & \dots \\ \psi_1(\mathbf{r}_2) & \psi_2(\mathbf{r}_2) & \dots & \psi_N(\mathbf{r}_2) \\ \psi_1(\mathbf{r}_p) & \psi_2(\mathbf{r}_p) & \dots & \psi_N(\mathbf{r}_p) \\ \psi_1(\mathbf{r}_p) & \psi_2(\mathbf{r}_p) & \dots & \psi_N(\mathbf{r}_p) \\ \psi_1(\mathbf{r}_p) & \psi_2(\mathbf{r}_p) & \dots & \psi_N(\mathbf{r}_p) \\ \dots & \dots & \dots & \dots \\ \psi_1(\mathbf{r}_p) & \psi_2(\mathbf{r}_p) & \dots & \psi_N(\mathbf{r}_p) \end{pmatrix} \begin{pmatrix} A_1 \\ A_2 \\ A_3 \\ \dots \\ A_N \end{pmatrix}$$

Where:

p = number of locations

N_t = number of time samples

N = number of modes

Fitting and the Pseudo-inverse

- Classic fitting problem for a data set: (x_i, y_i)

$$\vec{y} = A \vec{x}$$

$$\vec{x} = A^{-1} \vec{y}$$

- $\vec{x} = (A^T A)^{-1} A^T \vec{y}$ if A is rectangular

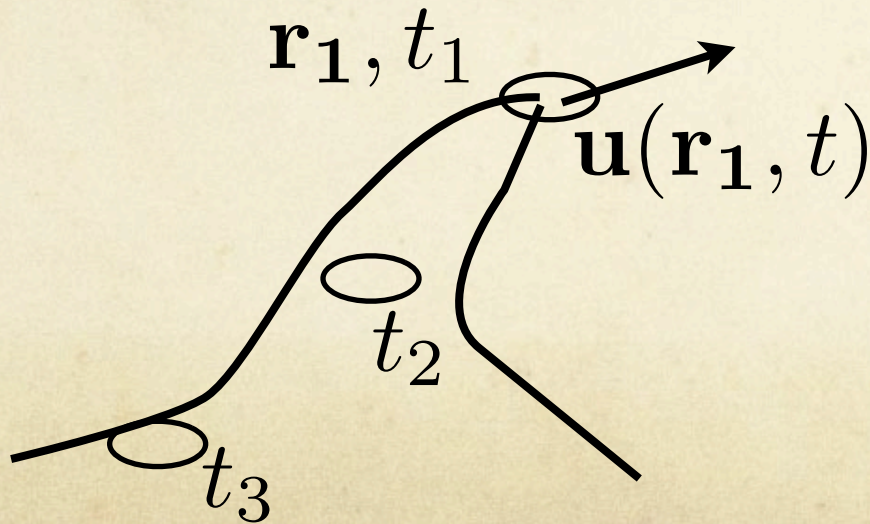
- Rows and columns cannot have zero or repeated entries

- Poor condition number for A leads to unstable solution, \vec{x}

- $\delta \vec{x} = A^{-1} \delta \vec{y}$ where small variations in y lead to large variations in x

Backing out the Time Dependence

- For currents at \mathbf{r}_1 , where is the current at t_2 ?
- Define a new position for each time by successively “backing out” the current.
- Prevents rows from repeating.
- Lowers the condition number of A.



Monte Carlo

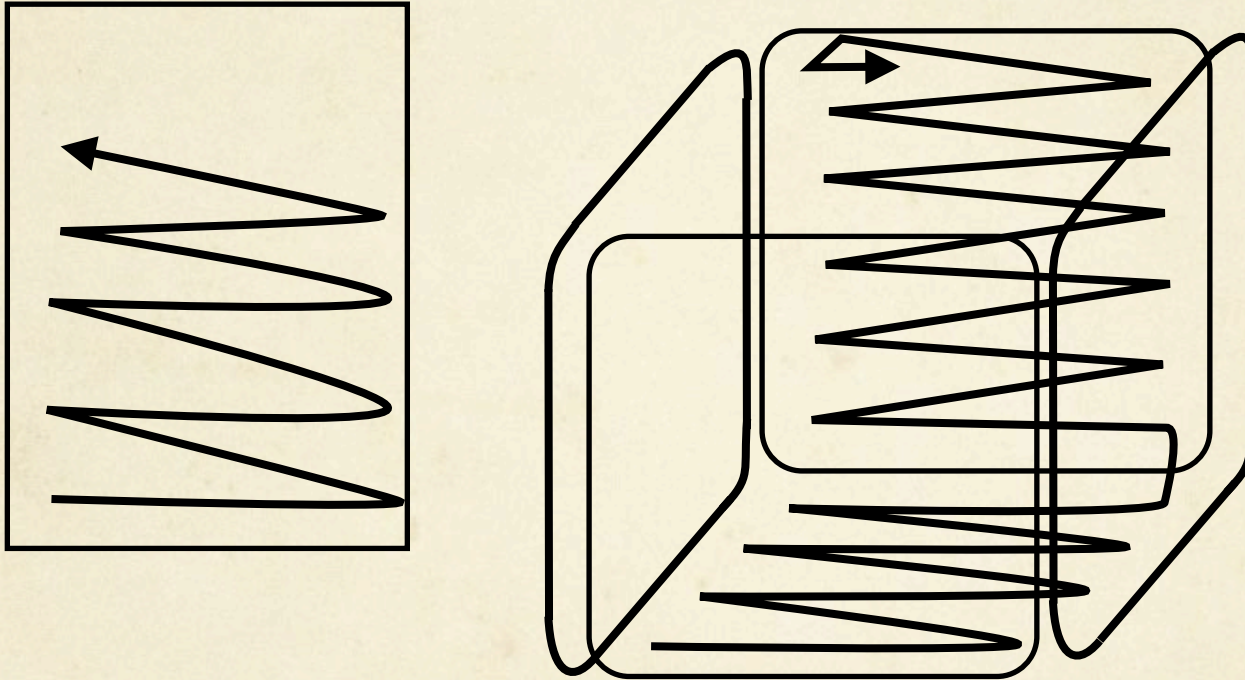
- By adding a slight variation to either \vec{y} or A , force errors in \vec{x} .
- Repeat the process 1000+ times and average the answer.

Power Spectra over Time

- Given a set of constant amplitudes for each time window, over all windows, assign a time dependence to each mode, $A_1(t)$...
- Re-adjust the modal matrix at each position based on the extracted time dependence for each mode.
- Repeat entire process until the time dependence for each mode becomes stable.

Data State Vector

- Whether 2D or 3D, state vector of the data is re-formatted as a column vector (rasterized).



- Number of populated terms in the state vector is always very sparse (typically 1%).

Conclusions

- Realistic Normal Mode Analysis must utilize sparse data over space.
- NMA provides an alphabet for how to discuss the power spectrum for a system - not a physical gyre representation.
- NMA - when successful provides full domain coverage given a limited number of spatially sampled locations.