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# Quantum thermodynamics for a model of an expanding universe

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# Relativistic and Quantum Physics

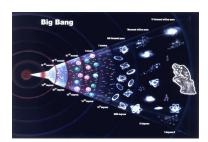


 ${\sf Relativity} + {\sf Quantum\ Information} = {\sf Relativistic\ Quantum\ Information}$ 

## Important achievements in general relativistic scenarios

- Three laws of black holes;
- Cosmology and big bang;
- Firewall issues.





#### Outlook

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- Quantum processes in the universe do not necessary involve large numbers of constituents.
- Important: Von N. Entropy of the Universe cannot change.

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Use quantum thermodynamics to understand work, entropy and energy flows in relativistic and cosmological setups.

#### Motivations

- The Universe is relativistic and quantum system and processes can involve small numbers of constituents.
- Mainly: We cannot compute energy, entropy and work flows in relativistic quantum systems.

## Expanding universe line/metric element

$$ds^2 = \Omega^2(\tau) \left[ -d\tau^2 + dx^2 + \ldots \right], \qquad g_{\mu\nu} = \Omega^2(\tau) \left( -1, 1, \ldots \right)$$

# Simple cosmology model

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## Scalar quantum field

$$\phi(t,x) = \oint_{t} \left[ u_{k} a_{k} + u_{k}^{*} a_{k}^{\dagger} \right], \quad \left[ a_{k}, a_{k'}^{\dagger} \right] = \delta^{d}(k - k')$$

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## Frequency shift and choice of conformal factor

$$\omega_{\mathsf{in}/\mathsf{out}} = \sqrt{k^2 + m\,\Omega^2(\tau_{\mp\infty})}, \qquad \Omega(\tau) = \sqrt{1 + \epsilon\,\big(1 + \mathsf{tanh}\big(\sigma\,\tau\big)\big)}$$

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## Bogoliubov transformations and squeezing

$$a_{\mathrm{out},k} = \cosh r_k \, a_{\mathrm{in},k} + e^{i\,\theta_k} \, \sinh r_k \, a_{\mathrm{in},-k}^\dagger, \qquad \tanh r_k = \frac{\sinh \left(\pi \frac{\omega_{\mathrm{out}} - \omega_{\mathrm{in}}}{2\sigma}\right)}{\sinh \left(\pi \frac{\omega_{\mathrm{out}} + \omega_{\mathrm{in}}}{2\sigma}\right)}$$

# Work and energy of two mode squeezing

All couple of modes (k, -k) decouple. We can focus on a single couple and redefine  $a_{\text{in},k} \equiv a_{\text{in}}$  and  $a_{\text{in},-k} \equiv b_{\text{in}}$ .

#### Initial Hamiltonian

$$H_{\rm in} = \omega_{\rm in} \left[ a_{\rm in}^{\dagger} a_{\rm in} + b_{\rm in}^{\dagger} b_{\rm in} + \frac{1}{2} \right]$$

#### Final Hamiltonian

$$H_{\mathrm{out}} = \omega_{\mathrm{out}} \left[ a_{\mathrm{out}}^{\dagger} a_{\mathrm{out}} + b_{\mathrm{out}}^{\dagger} b_{\mathrm{out}} + \frac{1}{2} \right]$$

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Start with an initial *thermal* state  $\rho$  with  $n_i$  particles. The work W done by spacetime is

## Work performed

$$W = \text{Tr} ((H_{\text{out}} - H_{\text{in}}) \rho)$$
  
=  $\omega_{\text{out}} n_c + (\omega_{\text{out}} - \omega_{\text{in}}) (n_i + 1)$ 

with  $n_c$  particles that are created.

## The "inner friction"

An adiabatic process would lead to a work cost  $W_{\mathrm{ad}}$  of the form

#### Adiabatic work

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## Our last step

We proceed to show that  $W_{\text{fric}}$  can be interpreted as an entropic quantity

## Forward process

$$p_{\rm in,out} = |\langle n_{\rm out} | n_{\rm in} \rangle|^2 \langle n_{\rm in} | \rho | n_{\rm in} \rangle$$

$$p_{\text{out,in}} = |\langle n_{\text{in}} | n_{\text{out}} \rangle|^2 \langle n_{\text{out}} | \rho | n_{\text{out}} \rangle$$

# Entropy and inner friction

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Fluctuation relation: 
$$s_{\text{in,out}} \coloneqq -\log(n_{\text{out}}|\rho|n_{\text{out}}) + \log(n_{\text{in}}|\rho|n_{\text{in}})$$

$$p_{\text{in,out}} = p_{\text{out,in}} e^{s_{\text{in,out}}}$$

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$$P_{\text{in,out}}(s) = \sum_{n_{\text{in}}, n_{\text{out}}} p_{\text{in,out}} \delta(s - s_{\text{in,out}})$$

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Average entropy: 
$$K(X||Y) := -\sum_n p_X(n) [\log p_Y(n) - \log p_X(n)] \ge 0$$

$$s = K(P_{\text{in,out}}||P_{\text{out,in}})$$

## Initial state

$$\rho = \frac{e^{-\beta_{in} H_{in}}}{\mathcal{Z}}$$

## Entropy

$$s_{\rm in,out} = \frac{\omega_{\rm in}}{T} n_c$$

## Results

## Initial state

$$\rho = \frac{e^{-\beta_{in} H_{in}}}{\mathcal{Z}}$$

## Entropy

$$s_{\rm in,out} = \frac{\omega_{\rm in}}{T} n_c$$

## Final result

$$s = \frac{\omega_{\rm in}}{T} n_c$$

This is our main result.

#### Extendable result

Number of created particles  $n_c = \sinh^2 r$ . This result can be extended to:

- \* Unruh effect:  $\tanh r = \exp\left[\frac{\hbar \omega}{k_B T_U}\right]$ ;
- \* Schwarschild black hole:  $\tanh r = \exp\left[\frac{\hbar \omega}{k_B T_H}\right]$ ;
- \* Analogue gravity models.

## Conclusions and outlook

#### Conclusions

- \* We have studied applications of quantum thermodynamics to setups that appear in quantum field theory;
- \* Have found some entropic quantity that increases in (simple) cosmological processes;
- \* The results apply to different cosmological and quantum field theoretical scenarios.

#### Outlook

- \* Can teach us more about the physics at the overlap of relativity and quantum mechanics;
- \* Drive future theoretical efforts to uncover novel physics;
- \* Hopefully: provide energy balance relations in quantum field theoretical scenarios.

# Thank You.



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